

4. Compressible flows

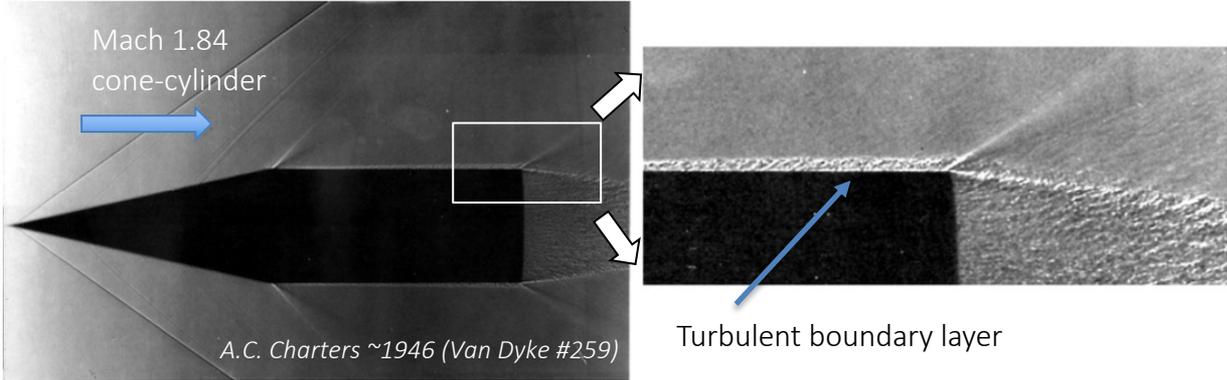
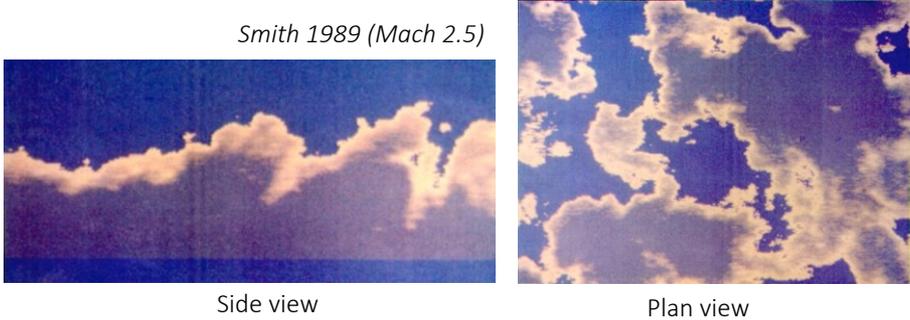
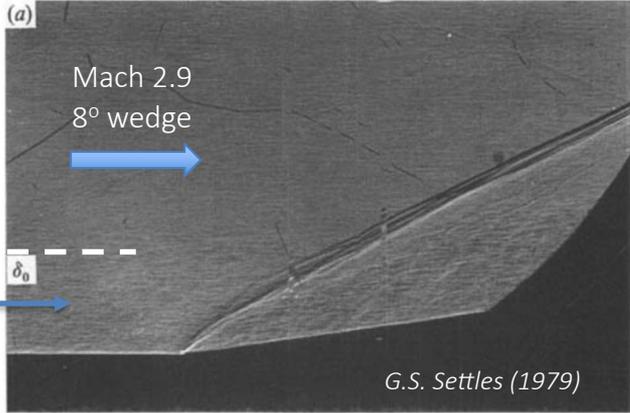
Alexander J. Smits

Department of Mechanical and Aerospace Engineering
Princeton University

Lecture 4, 25 August 2023
Les Houches School of Physics



High speed flows





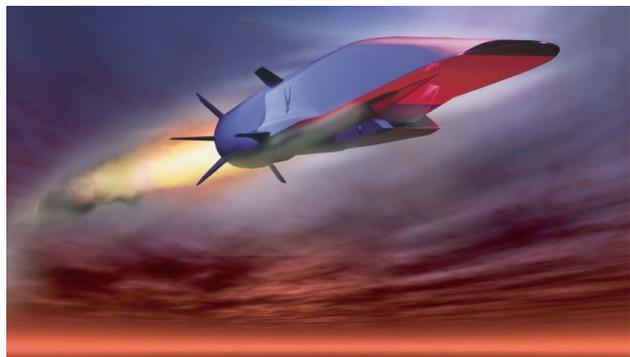
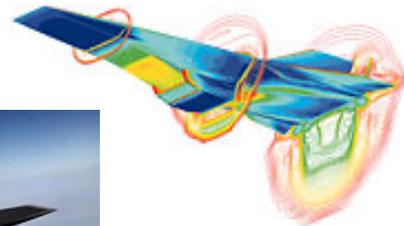
Space Shuttle (first flight 1981)



NASP X-30 (proposed 1986)



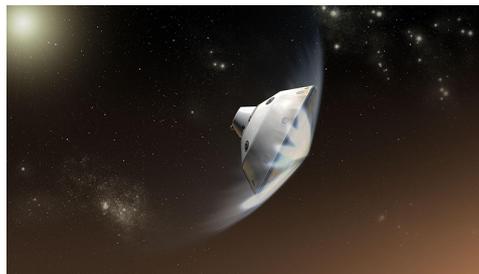
NASA X-43A/Hyper-X (first flight 2004)



Boeing X-51A Waverider (first flight 2010)



AF X-37B (first flight 2010)

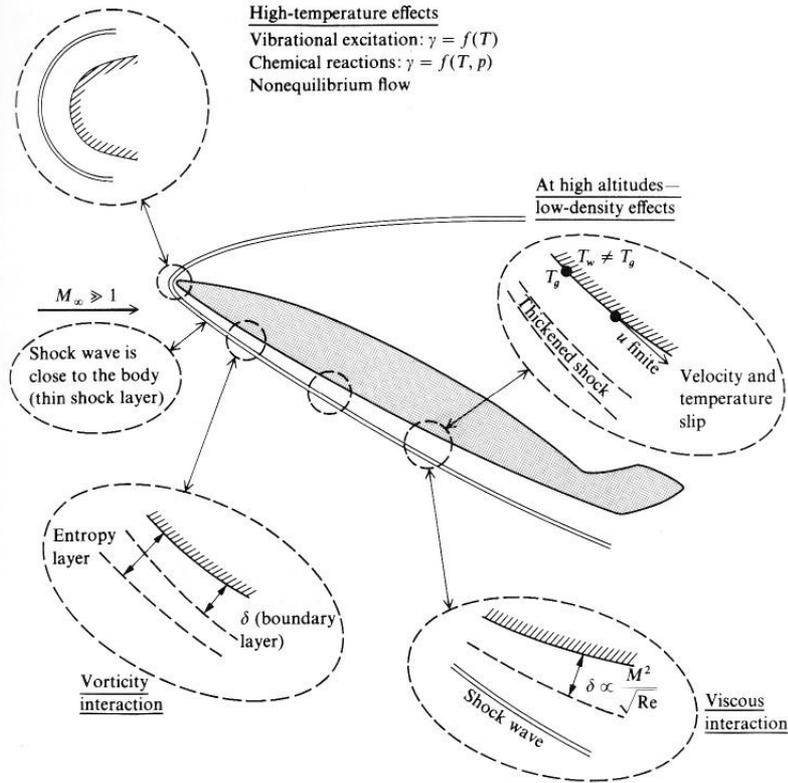


Curiosity Mars probe (2012)

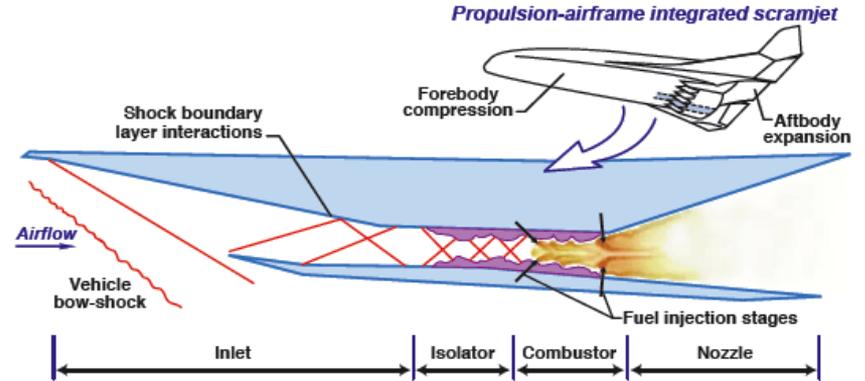


DARPA Glide Breaker

Hyperersonic vehicle flow fields



(Anderson, 1989)



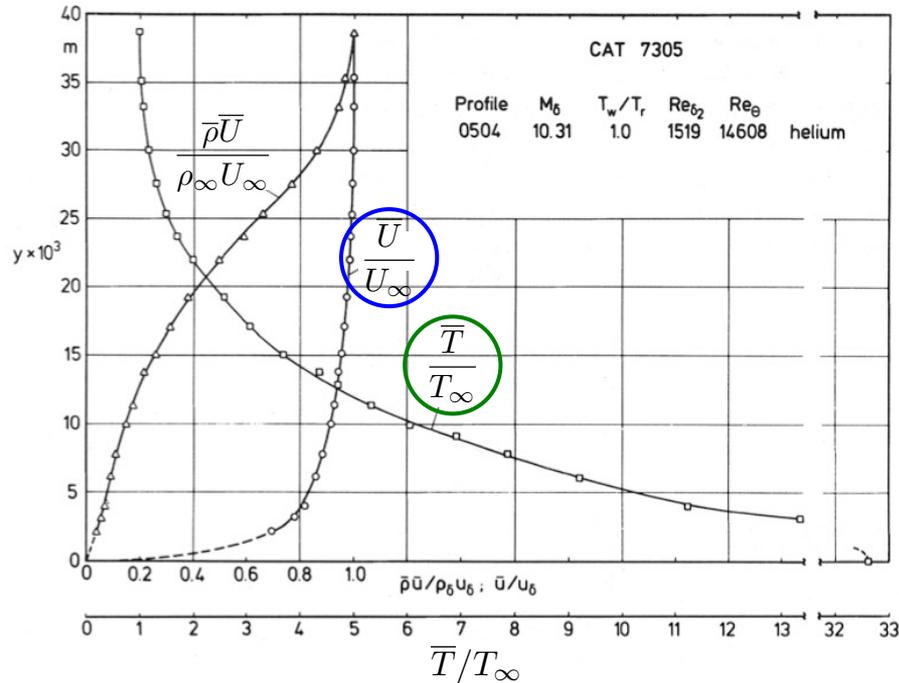
(NASA)



(Aerojet Rocketdyne)

Mean flow in high-speed boundary layers

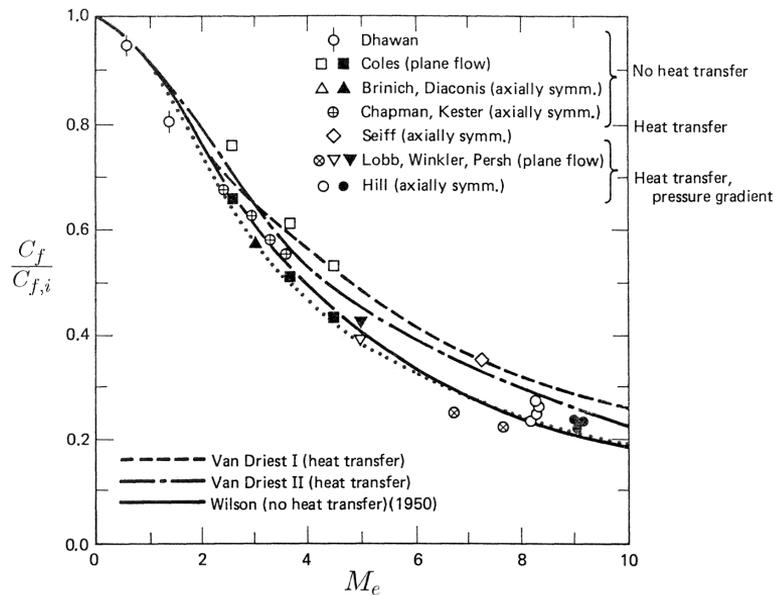
Ma = 10.3 (helium)



- Mean velocity profile looks similar to that in incompressible flow but scales differently
- Temperature increases near the wall
- Density decreases near the wall
- Need to account for density changes

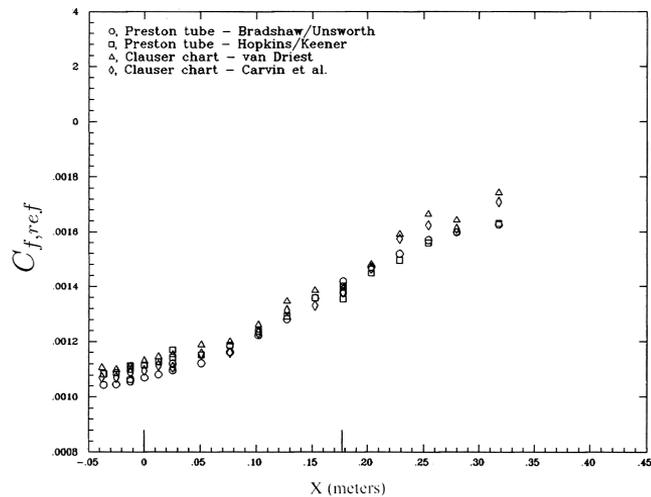
Skin friction

Zero pressure gradient



ρ_w decreases with Mach number

Adverse pressure gradient



ρ_w increases with compression

Compressible turbulent flows and Reynolds number

- In a compressible flow, we have a choice of Reynolds numbers
- Friction Reynolds number

$$Re_\tau = \frac{\delta}{\nu_w/u_\tau} = \frac{\delta u_\tau}{\nu_w}$$

- Momentum thickness Reynolds number (freestream viscosity or wall viscosity)

$$Re_\theta = \frac{U_e \theta}{\nu_e} \quad Re_{\theta_w} = \frac{U_e \theta}{\nu_w} \quad (Re_\theta > 200 Re_{\theta_w} \text{ for Mach 10 helium flow})$$

$$Re_{\delta_2} = \frac{\text{highest momentum flux}}{\text{highest shear stress}} = \frac{\rho_e U_e \theta}{\mu_w}$$

$$(Re_\theta > 10 Re_{\delta_2} \text{ for Mach 10 helium flow})$$

- The inner layer thickness relative to the boundary layer thickness increases with Ma

Compressible turbulent flows and time scales

- The friction Reynolds number is often interpreted as a ratio of length scales
- What about time scales?

- Outer time scale: $\frac{\delta}{U_\infty}$ Inner time scale: $\frac{\nu_w}{u_\tau^2}$

- Ratio outer to inner: $R_\tau = \frac{u_\tau}{U_\infty} \left(\frac{\delta u_\tau}{\nu_w} \right) = \sqrt{\frac{C_f}{2}} Re_\tau$

- Time scale ratio decreases with Mach number

What do we know or expect for the mean flow?

1. van Driest transformation collapses the mean velocity profile onto the incompressible form of the log law, and the incompressible form of the wake

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- In the overlap between the inner and outer scaling, we have a log-law

- Best expressed as

$$\frac{\partial \bar{U}}{\partial y} = \frac{\sqrt{\tau_w / \rho}}{\kappa y}$$

← velocity scale
← length scale

- Hence:

$$U_{VD} = \int_{U_1}^U \sqrt{\frac{T_w}{T}} dU = \int_{U_1}^U d\left(\frac{U}{u^*}\right)$$

where $u^* = u_\tau \sqrt{(\rho_w / \rho)}$

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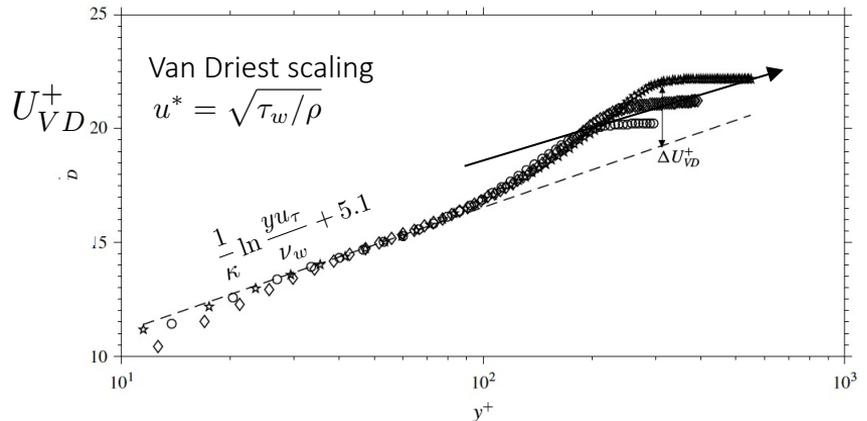
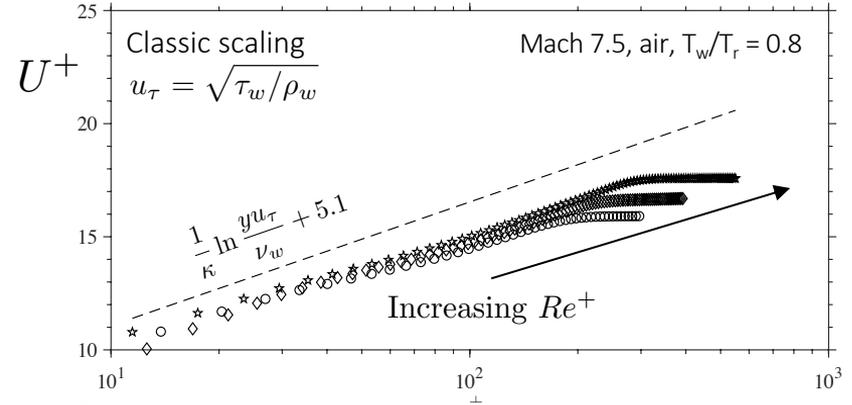
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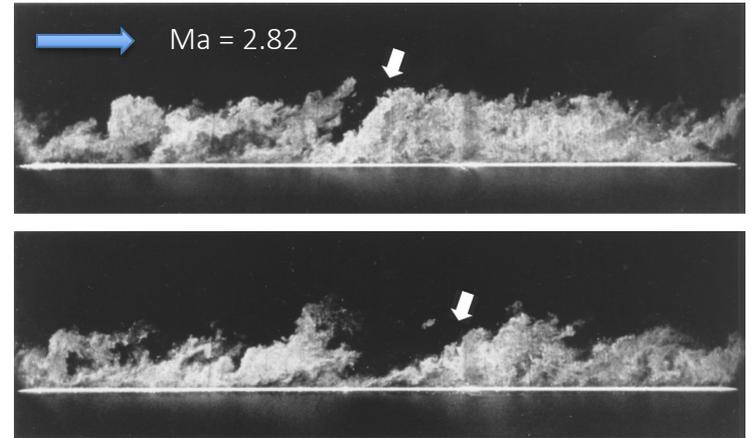
(See also Lee, Helm, Martin, Williams 2023)



What do we know or expect for the turbulence?

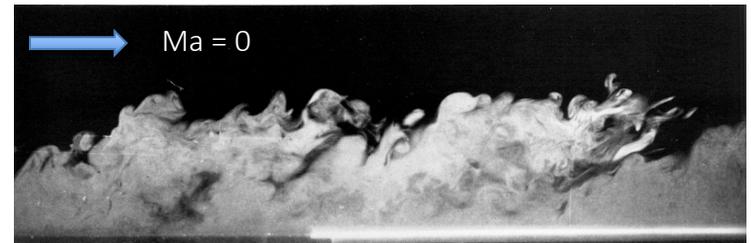
2. Mach number effects act only through the mean density (temperature) variation, and effects due to density fluctuations (shocklets, for example) are negligible even at high Mach number

Mie scattering from acetone droplets



Smith (1989); Smith & Smits (1995)

Smoke flow visualization



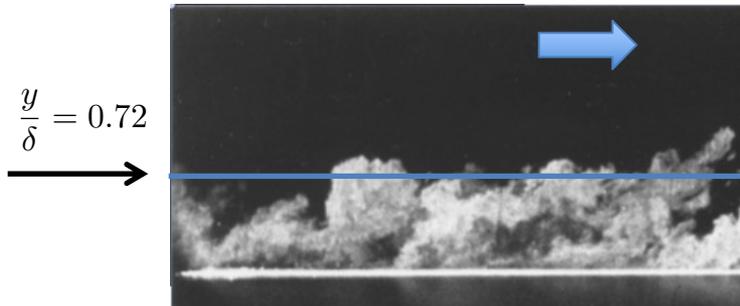
Falco (1977)

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Rayleigh scattering (FRS):

- Scattering from particles with $d < \lambda$
- ~ 10 nm clusters of H_2O , condensed in nozzle from naturally occurring water in air supply, or
- ~ 10 nm clusters of CO_2 , condensed in nozzle from CO_2 gas injected upstream (Miles & Forkey 1991)



Mach 2.9, $Re_\theta = 78,000$



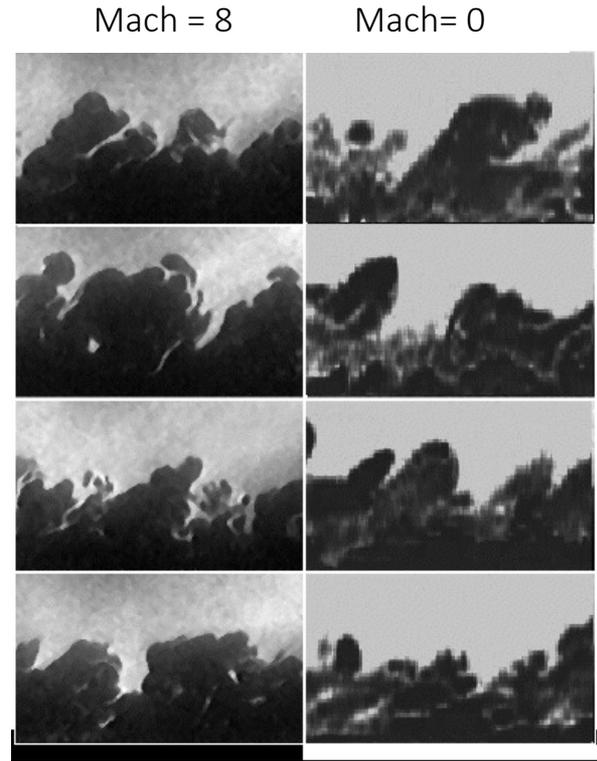
Side view



Plan view

Scaling the turbulence

- 3. Morkovin scaling collapses the turbulent stress profiles onto the incompressible data, at least for the overlap region and the outer flow



FRS – Baumgartner et al. (1997) ; PLIF - Delo & Smits 1997)

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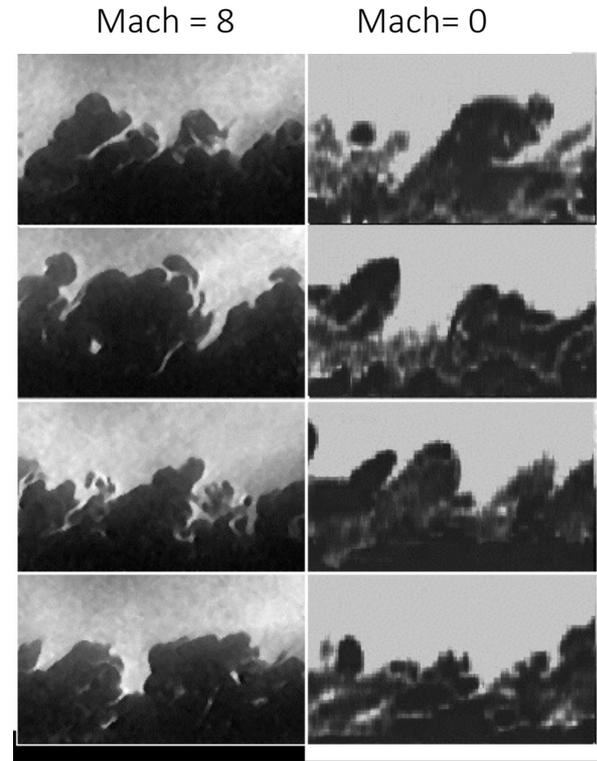
Morkovin's hypothesis (1961):

- “The essential dynamics of these shear flows will follow the incompressible pattern”

$$u_\tau = \sqrt{\tau_w / \rho_w}$$

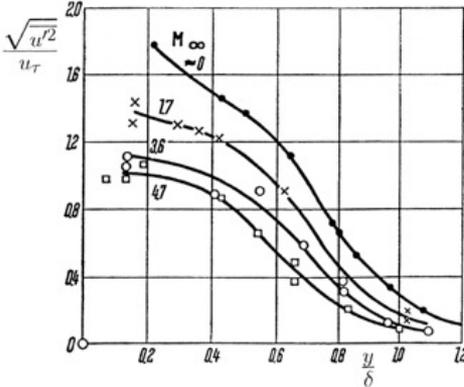


$$u^* = \sqrt{\tau_w / \rho}$$

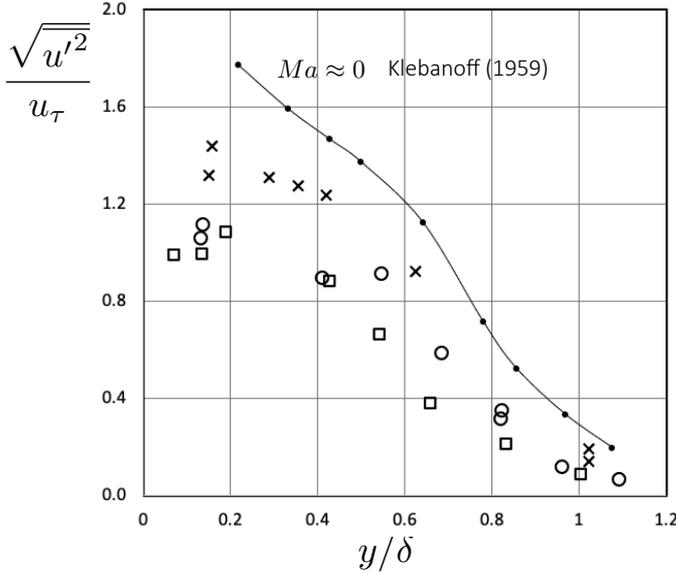


Scaling the turbulence

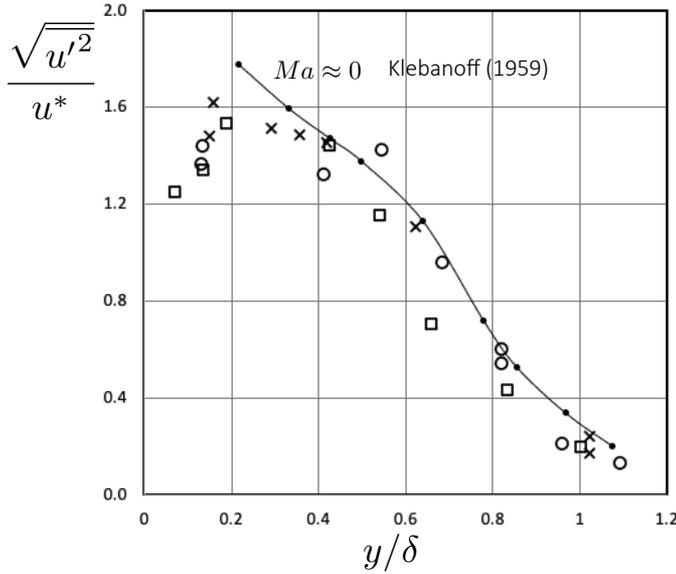
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Classic scaling



Morkovin scaling

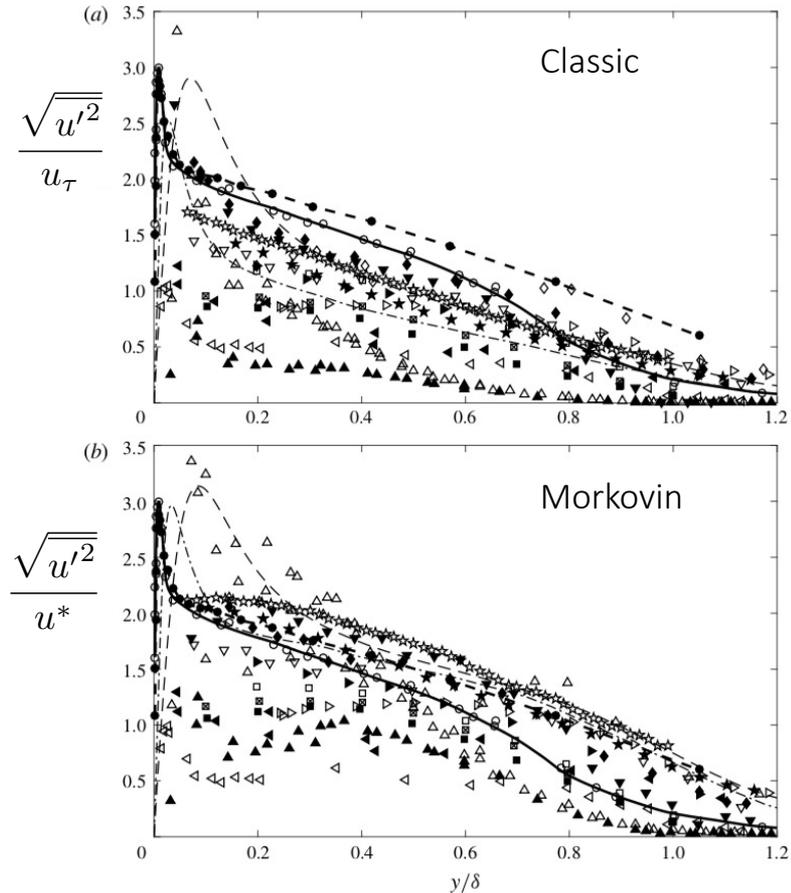


Historical record

Including all experimental data, up to $Ma = 11$,
obscures any conclusion

Previous hot-wire measurements often suffer from

- inadequate frequency response
- mixed-mode response/calibration
- inadequate spatial resolution.



*Williams,
Sahoo,
Baumgartner
& Smits
(2018)*

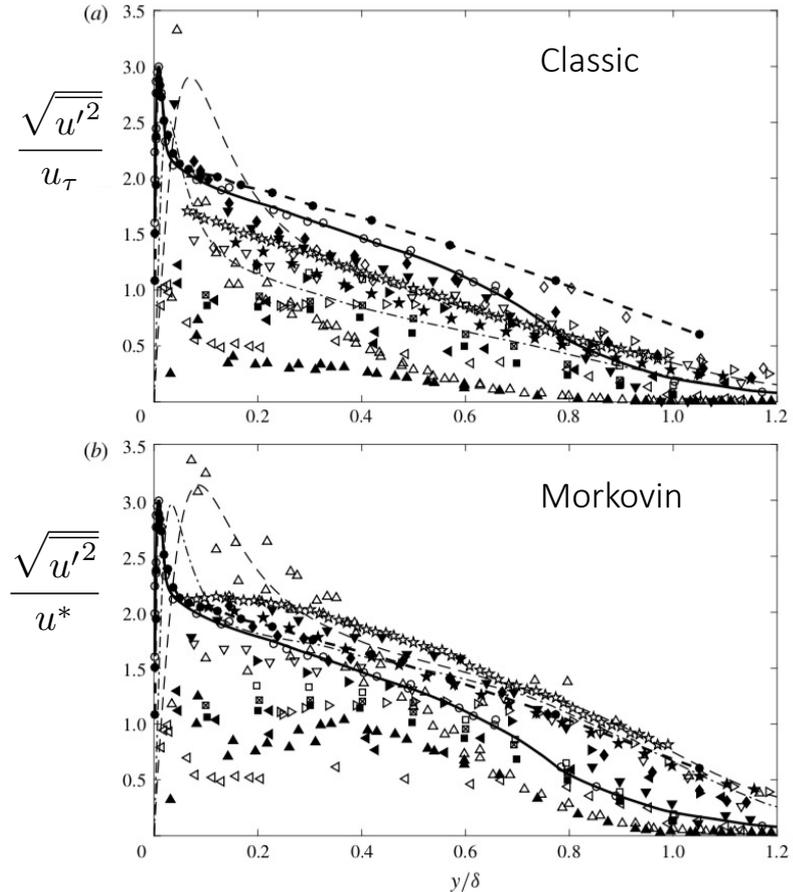
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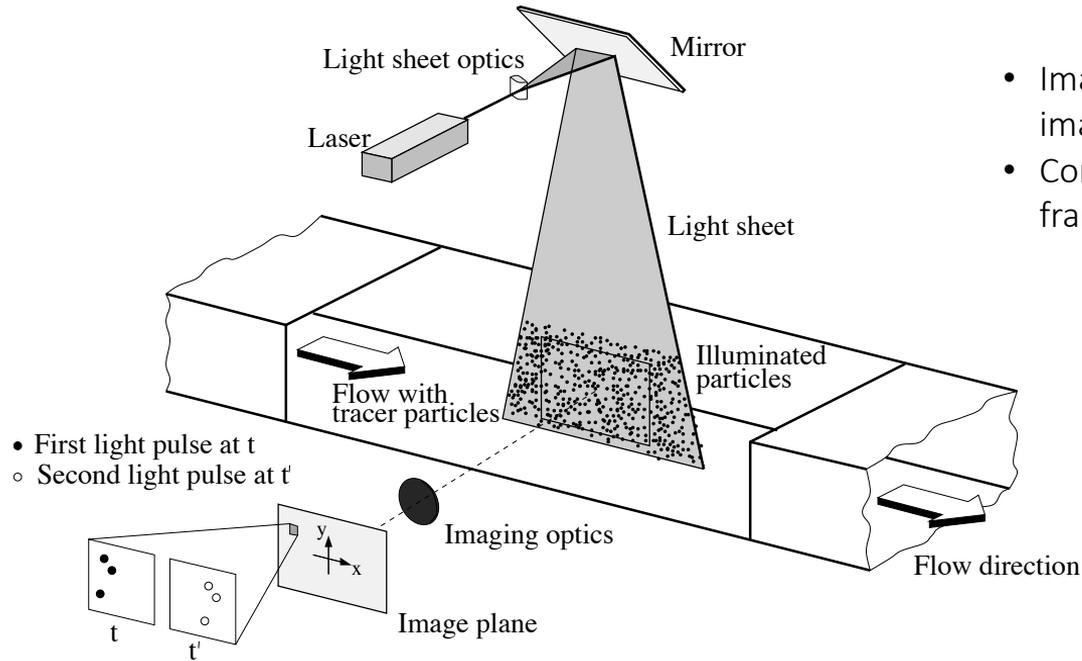
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What about PIV at high Mach number?



*Williams,
Sahoo,
Baumgartner
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(2018)*

Particle image velocimetry

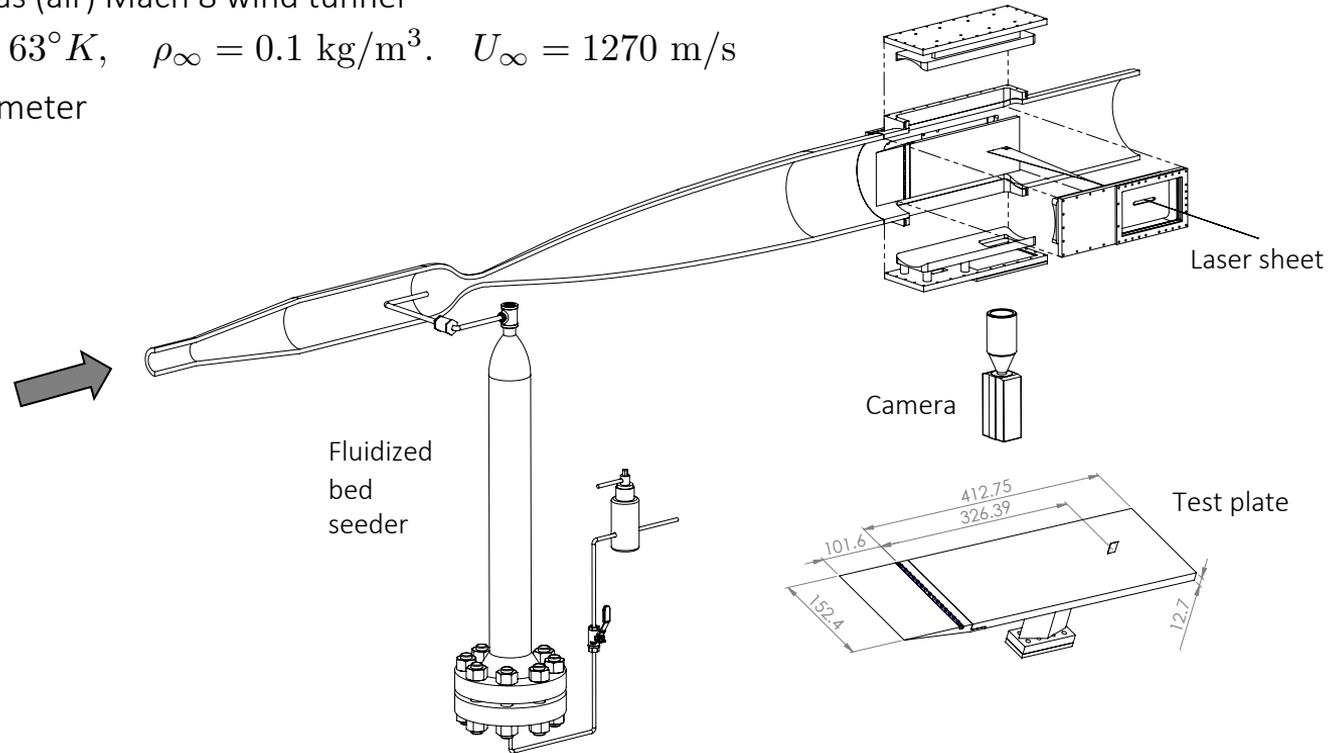


- Image particles using a laser and a fast, two-image camera
- Correlate motion of particles in successive frames

- Direct measure of velocity
- Provides spatial information
- Particles need to follow the flow

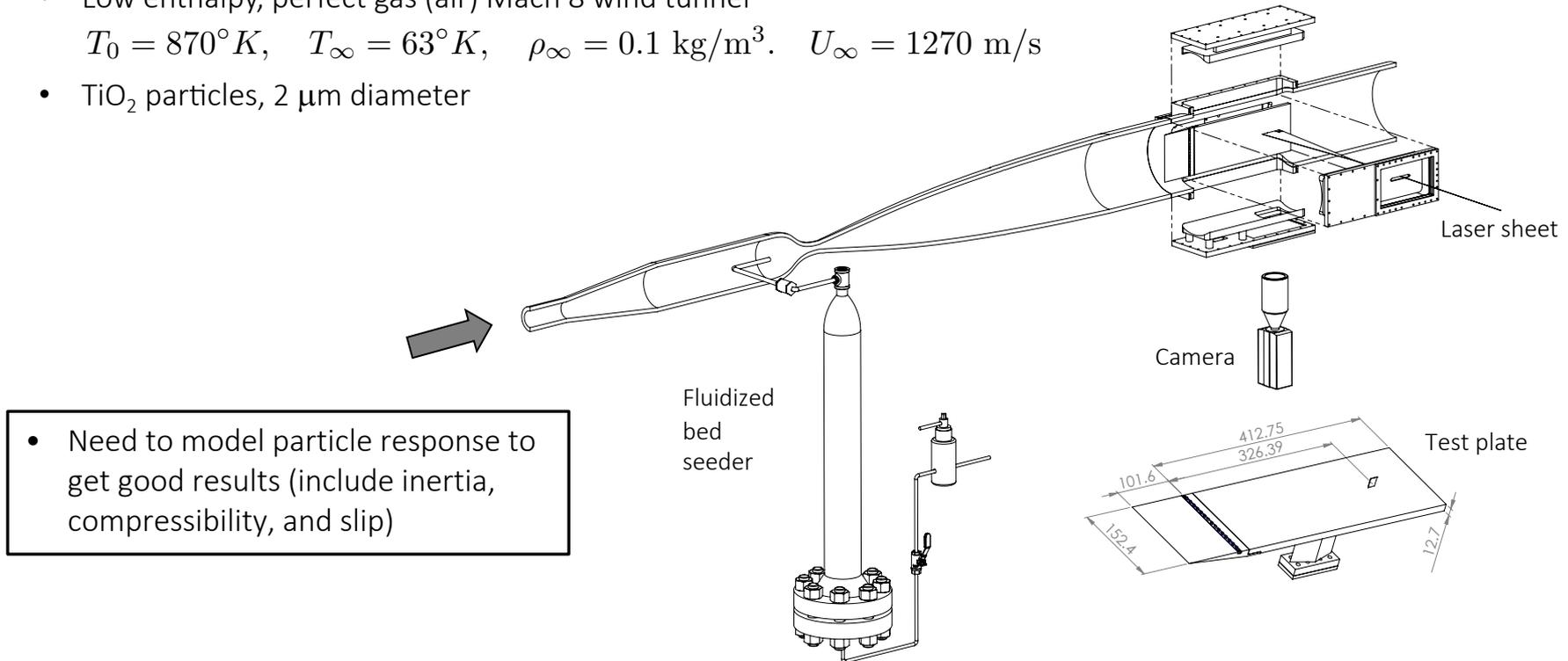
Mach 8 wind tunnel

- Low enthalpy, perfect gas (air) Mach 8 wind tunnel
 $T_0 = 870^\circ K$, $T_\infty = 63^\circ K$, $\rho_\infty = 0.1 \text{ kg/m}^3$. $U_\infty = 1270 \text{ m/s}$
- TiO_2 particles, $2 \mu\text{m}$ diameter



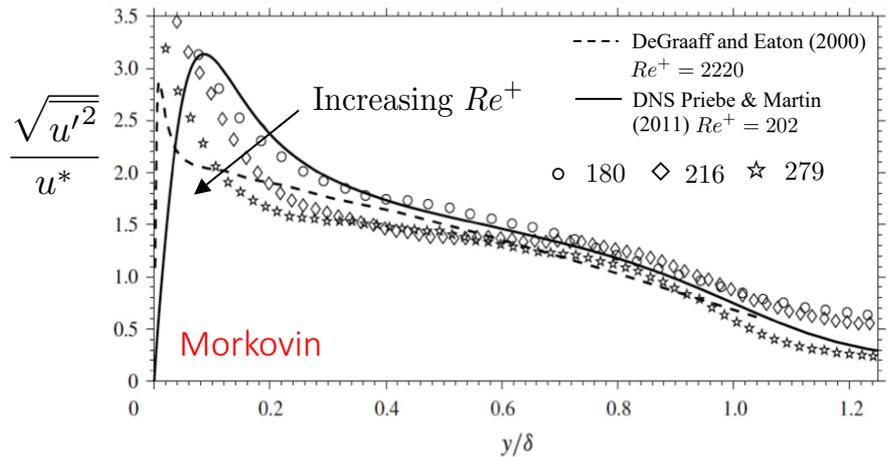
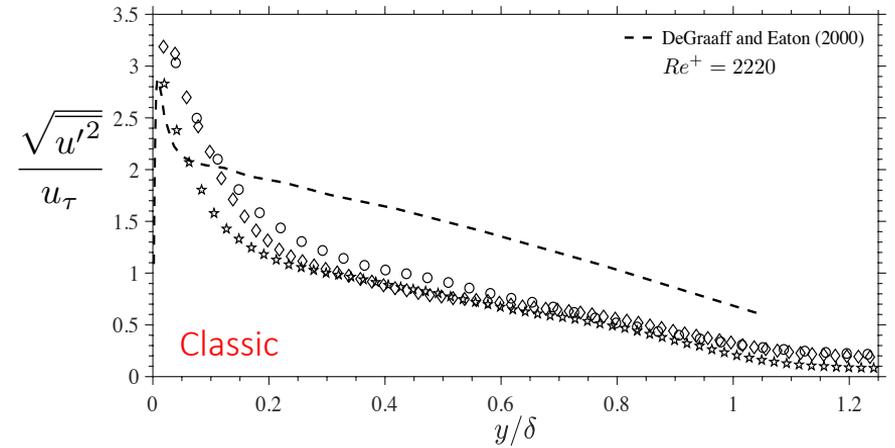
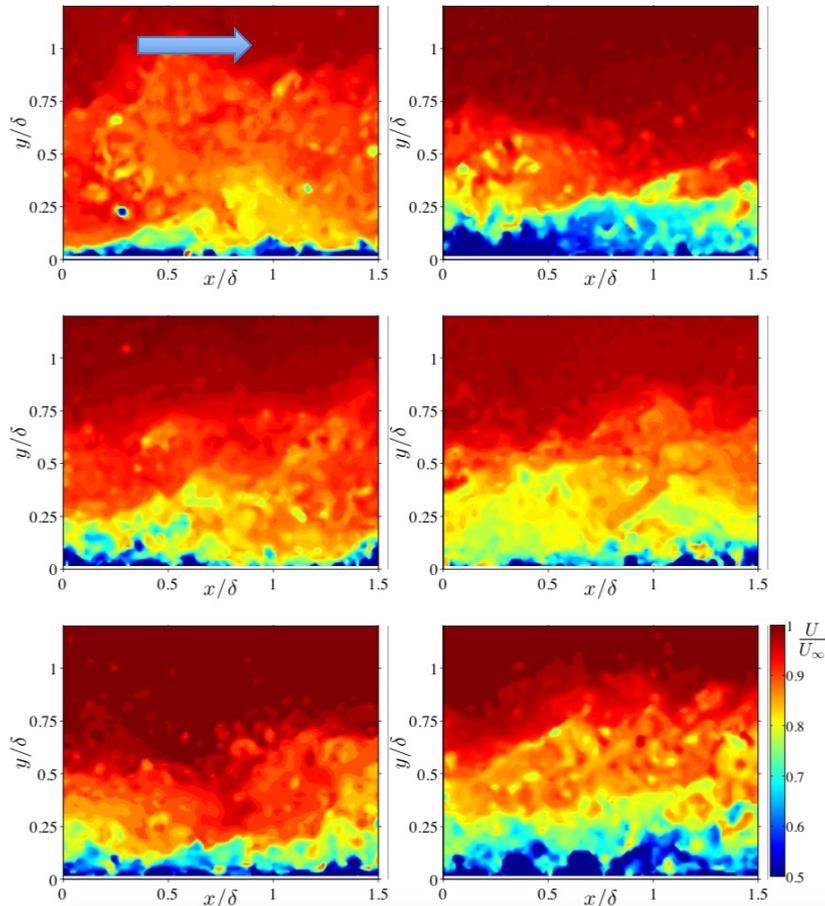
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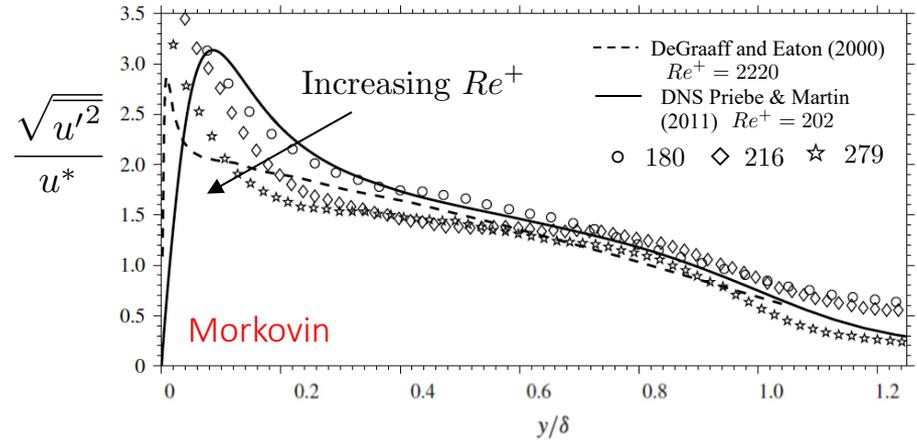
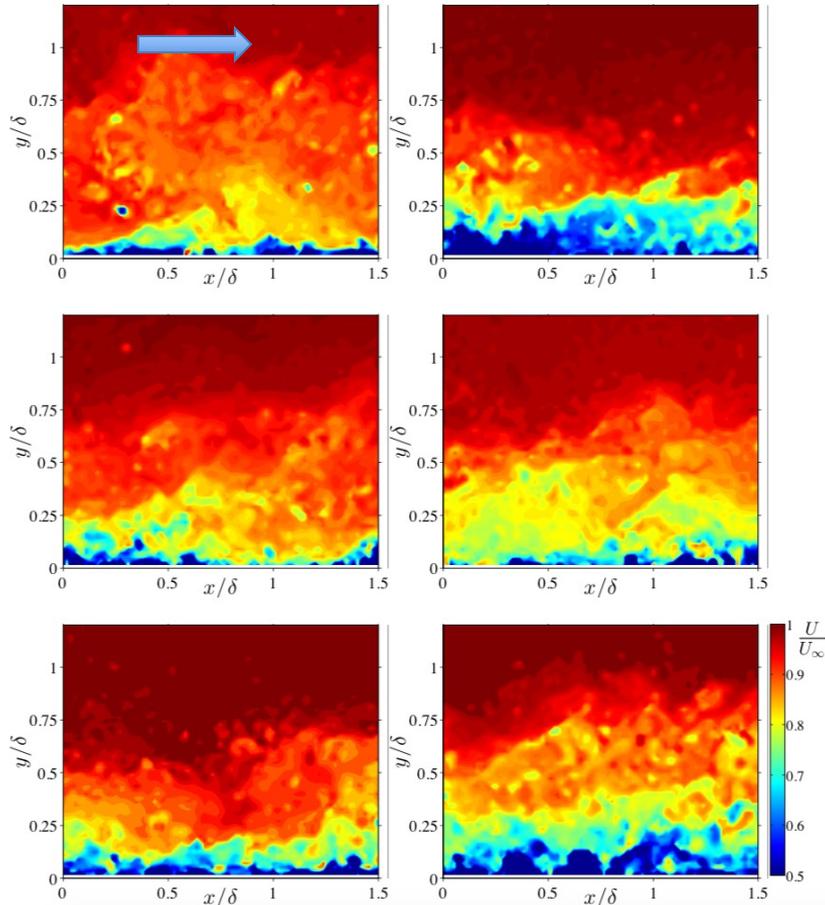


• Need to model particle response to get good results (include inertia, compressibility, and slip)

PIV results $Ma = 7.5$

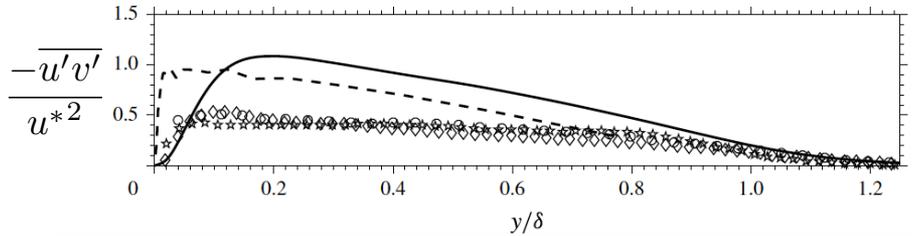
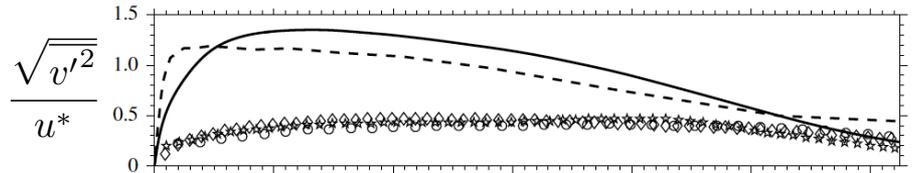
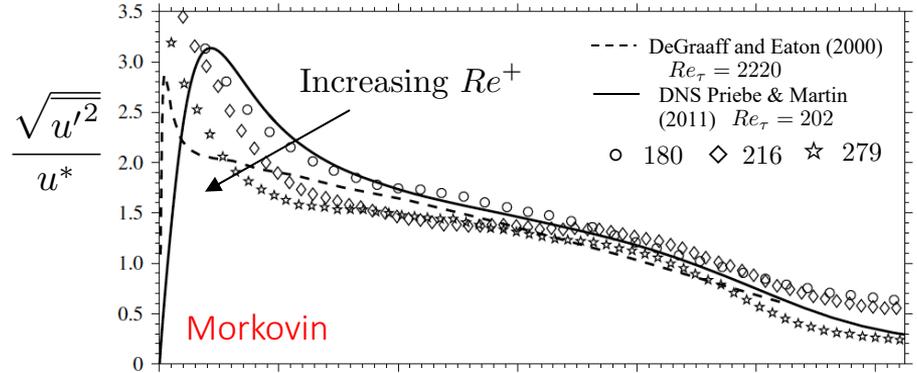
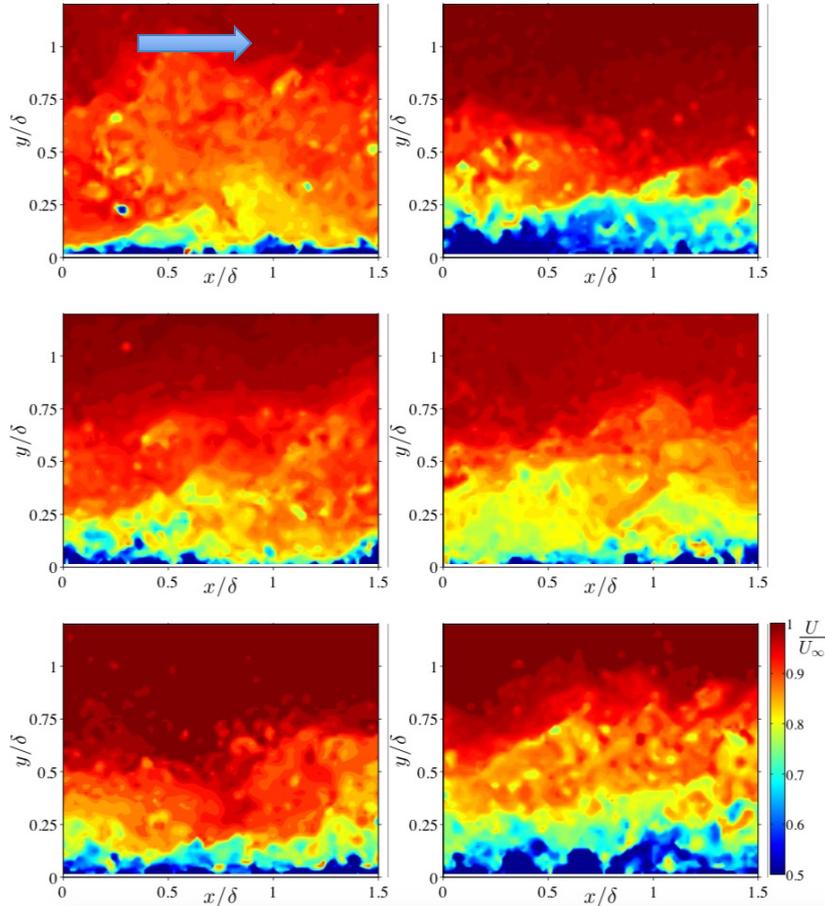


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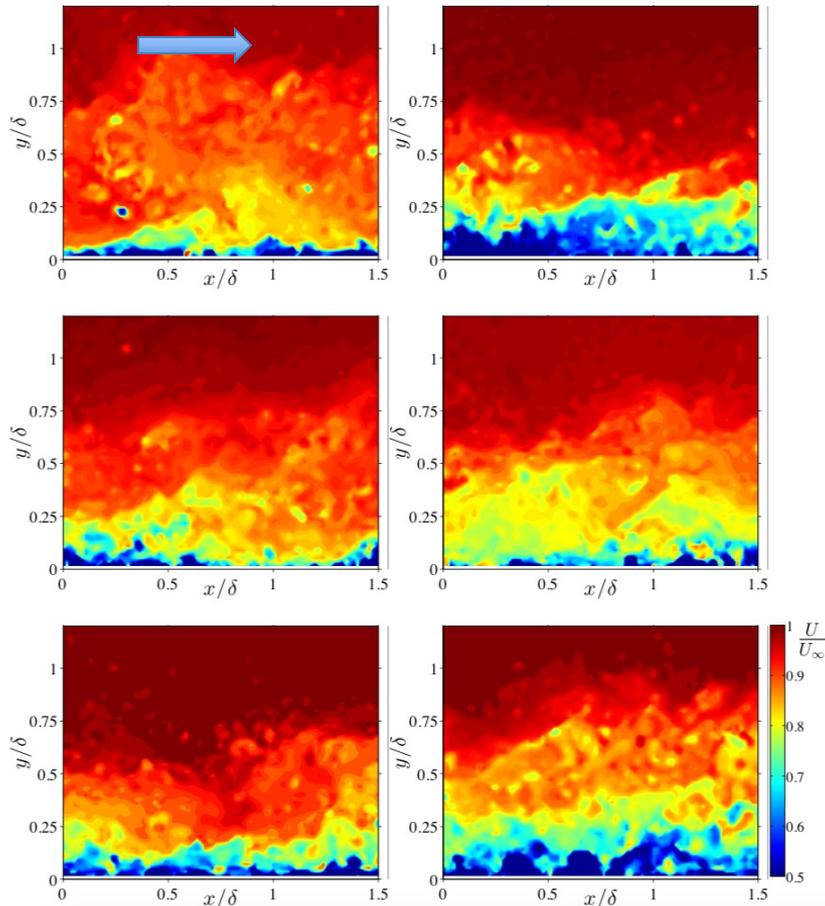


- Results agree well with DNS at the same Reynolds number
- Show the expected behavior with increasing Reynolds number
- Support Morkovin scaling
- But...

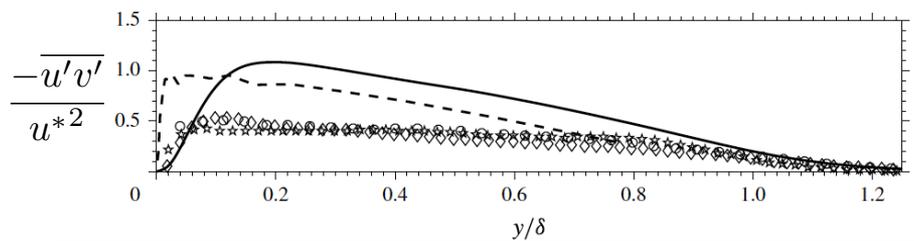
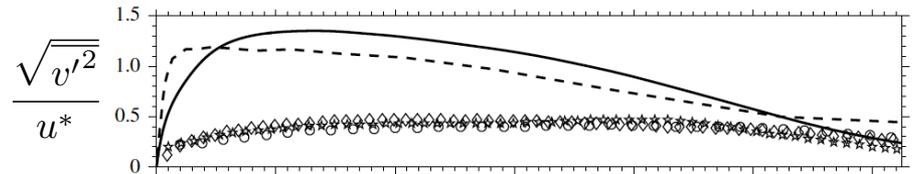
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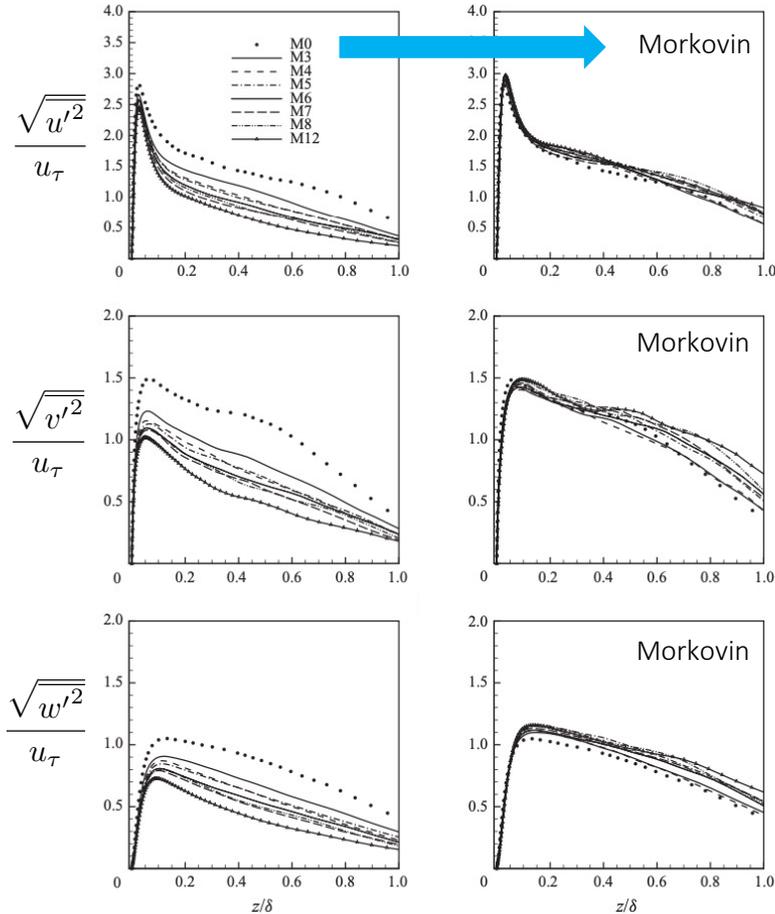
PIV results $Ma = 7.5$



- It is very difficult to measure the wall-normal component v'
- This is a general result for all experimental methods and almost all experiments



What about DNS?



- DNS of turbulent boundary layers
- (small box DNS, $\sim 10\delta$, bigger box results soon)
- Matched Reynolds numbers ($570 < Re^+ < 377$)
- $0.3 < Ma < 11.9$
- $1 < T_w/T_\infty < 28$

Profiles collapse in Morkovin scaling, for all three components, even at $Ma = 12$

Strong Reynolds Analogy between temperature and velocity

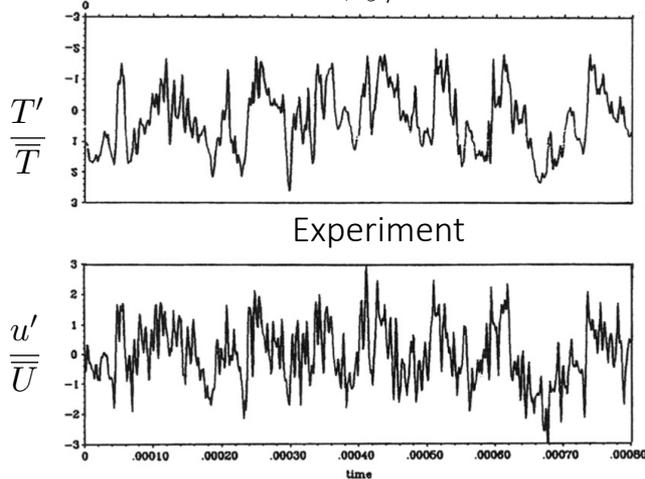
One-dimensional energy equation: $h_0 = h + \frac{1}{2}U^2$

Perfect gas with $T'_0 = 0 \rightarrow$

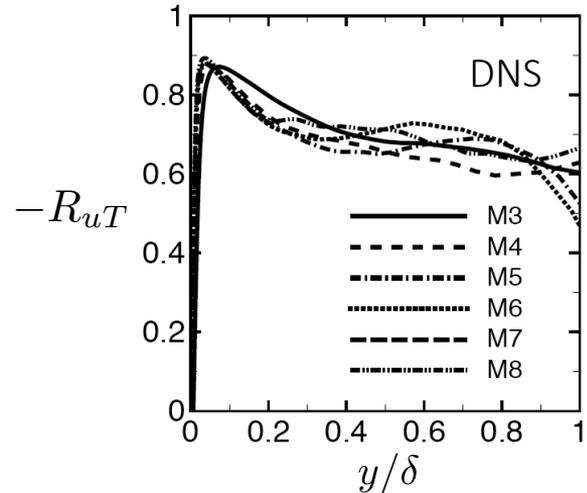
$$\frac{T_{rms}}{\bar{T}} = (\gamma - 1)M^2 \frac{u_{rms}}{\bar{U}},$$

$$R_{uT} = \frac{\overline{u'T'}}{u_{rms}T_{rms}} = -1$$

Mach 2.9, $y/\delta = 0.6$



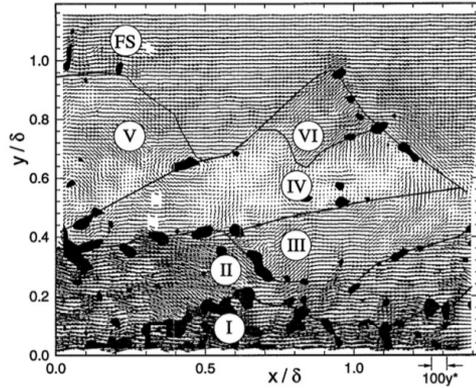
Smith & Smits (1993)



Wu & Martin (2004)

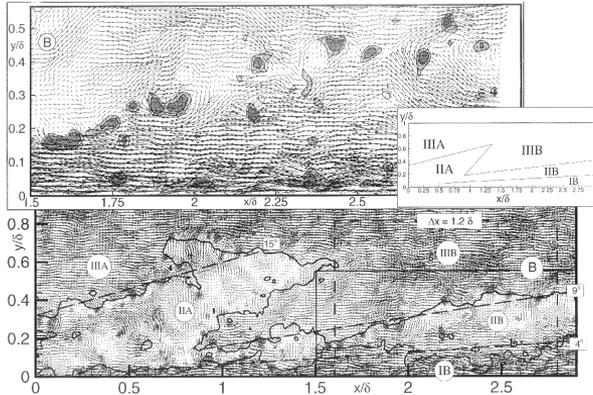
Uniform momentum zones (UMZ)

$Re^+ \approx 2600$



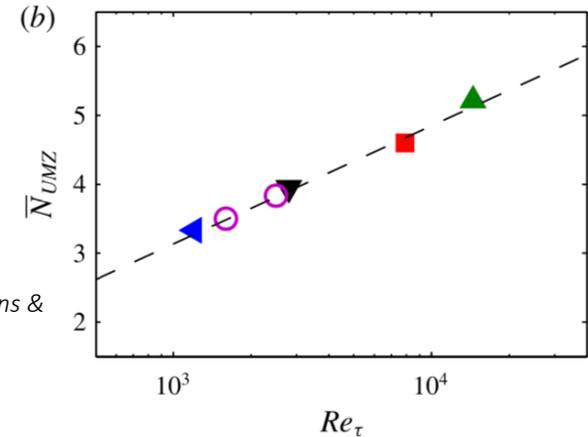
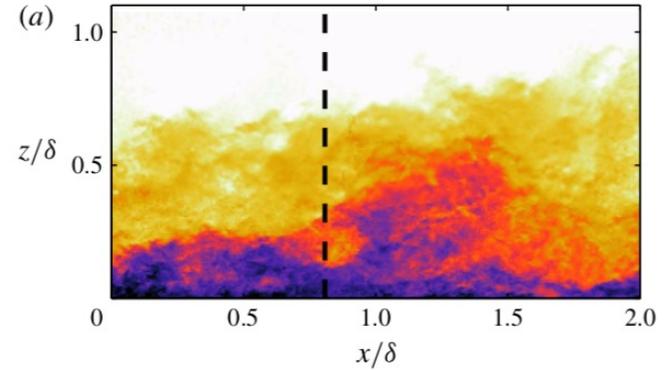
Meinhart & Adrian (1995)

- Number of UMZ's scales logarithmically with Re^+
- But not as fast as the number of hierarchies



Adrian, Meinhart & Tomkins (1999)

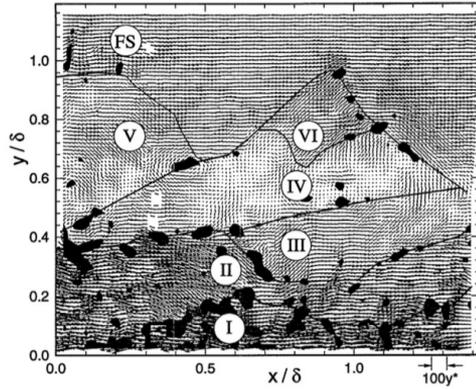
$Re^+ \approx 8000$



de Silva, Hutchins & Marusic (2016)

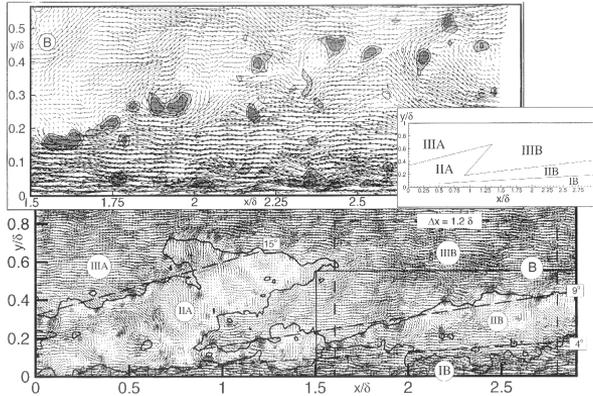
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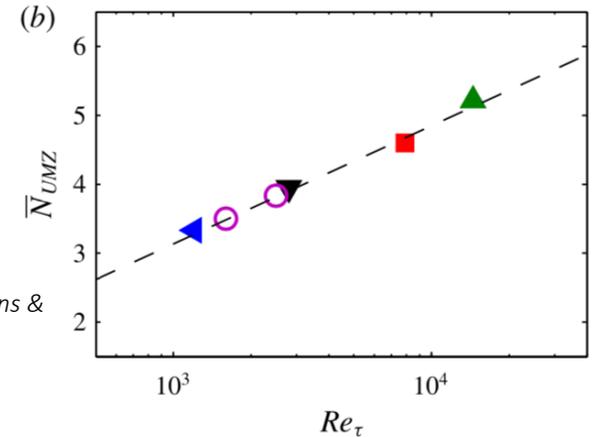
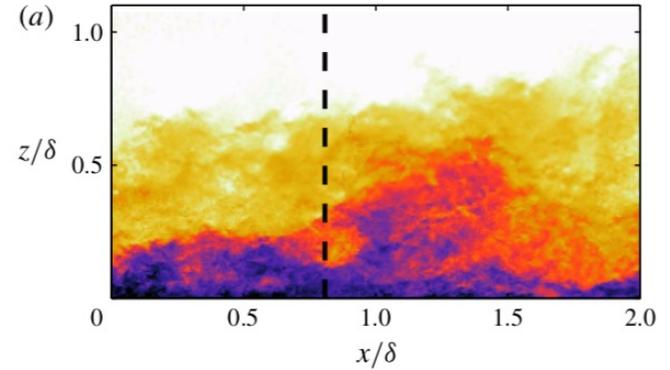
Meinhart &
Adrian (1995)

Incompressible flow
Mach = 0



Adrian, Meinhart & Tomkins (1999)

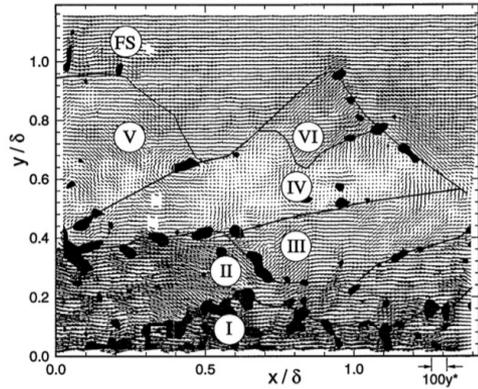
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de Silva, Hutchins &
Marusic (2016)

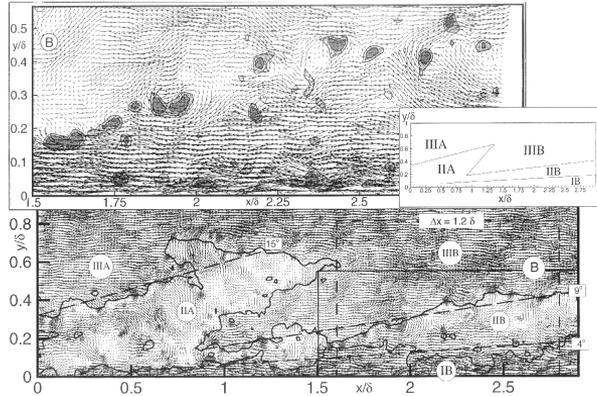
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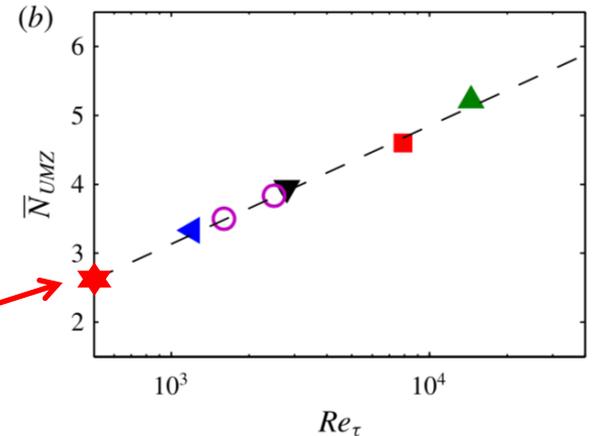
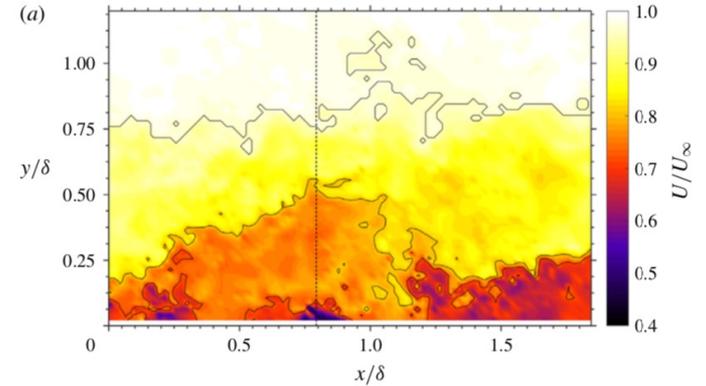
Meinhart & Adrian (1995)

Hypersonic flow
Mach 7.5



Adrian, Meinhart & Tomkins (1999)

$Re_\tau \approx 280$

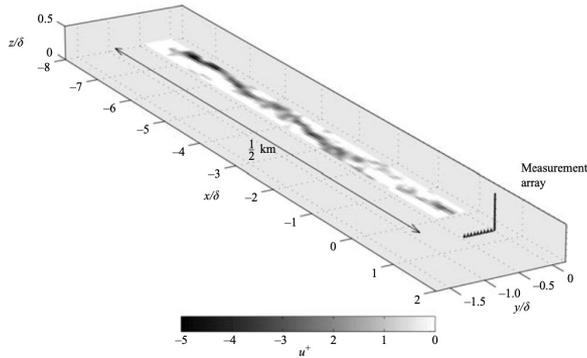


Mach 7.5

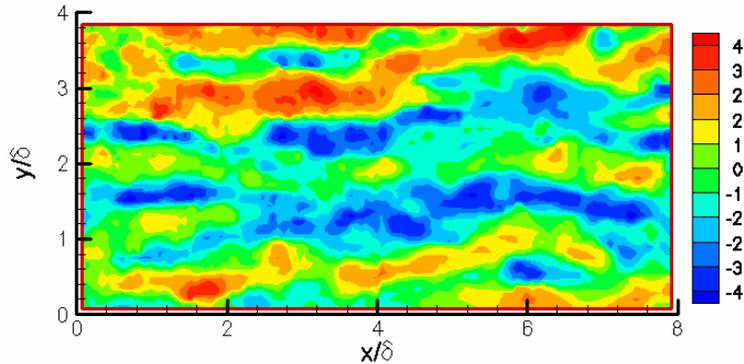
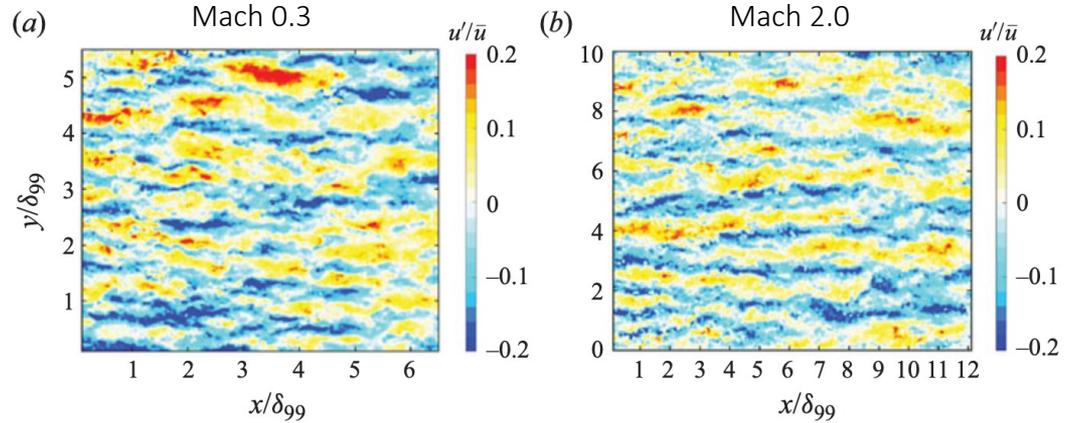
Williams, Sahoo, Baumgartner & Smits (2018)

Superstructures (VLSMs)

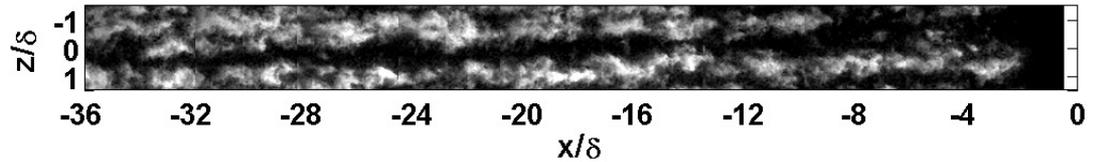
Atmospheric boundary layer
Hutchins & Marusic (2007)



Spanwise spacing $\approx \delta$ (may increase with Mach number (*Bross et al. 2021*))



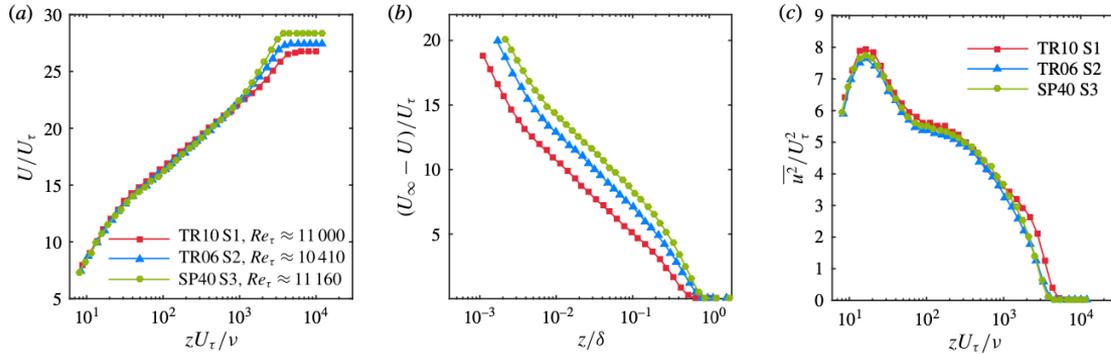
Mach 2
Ganapathisubramani et al. (2006)



Beyond flat plate boundary layers

- Van Driest and Morkovin work well for flat plate boundary layers
- What about “real” world effects?
 - Tripping effects
 - Transitional effects
 - Surface curvature
 - Görtler vortices
- What about complex flows?
 - Shock wave boundary layer interactions
 - Superstructures
 - Görtler vortices

Tripping effects

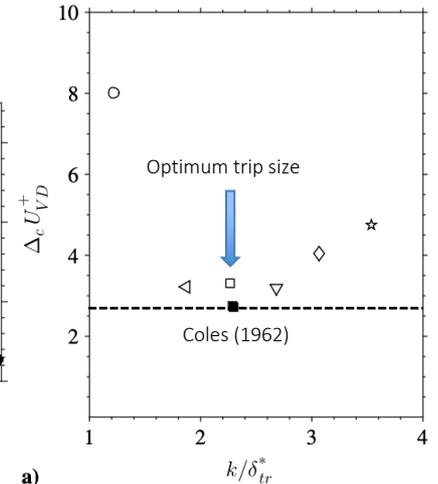
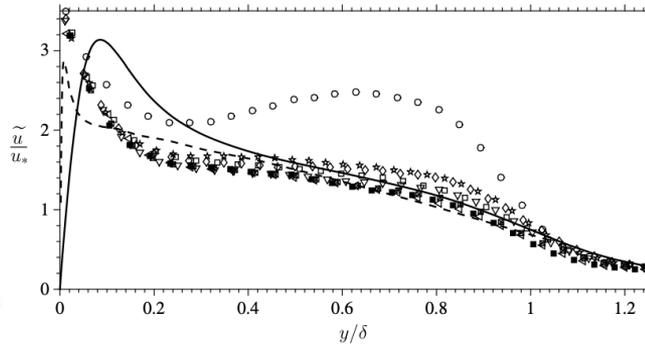
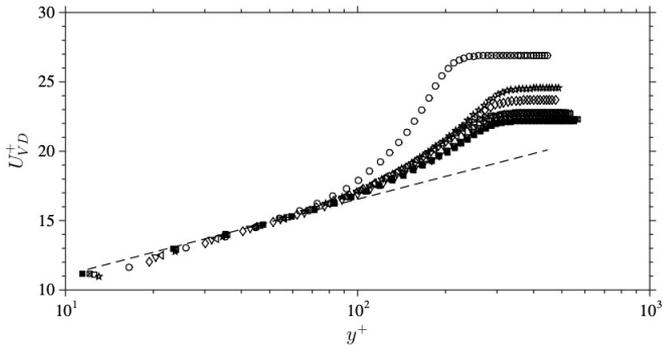


Subsonic flow
Marusic et al. (2015)

P40 grit sandpaper, 6mm
 and 10 mm threaded rod,
 matched $Re_\tau \approx 11000$

Mach 7.6

Williams & Smits (2017) Pin type stimulator trip

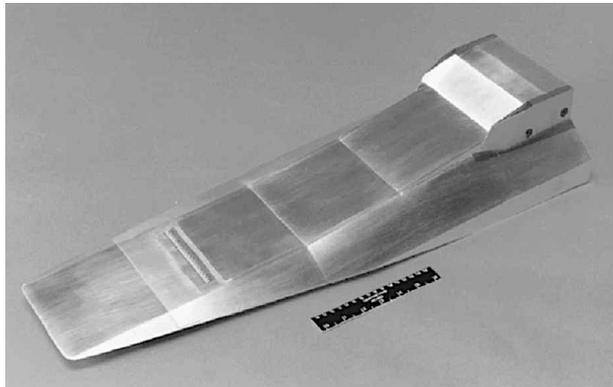


a)

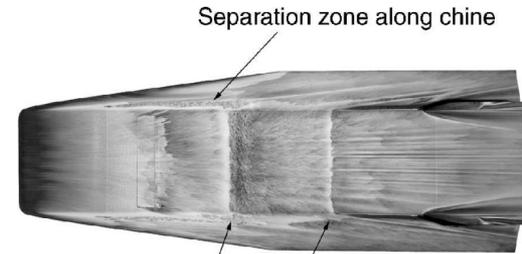
Tripping effects



- Hyper-X tripping at Mach 6, $\alpha = 2^\circ$

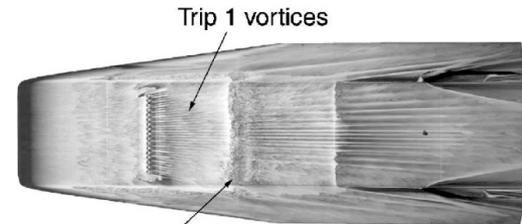


Berry et al. (2001)



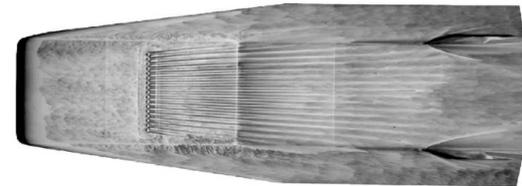
No trip

Separations at 1st and 2nd corner



Critical roughness

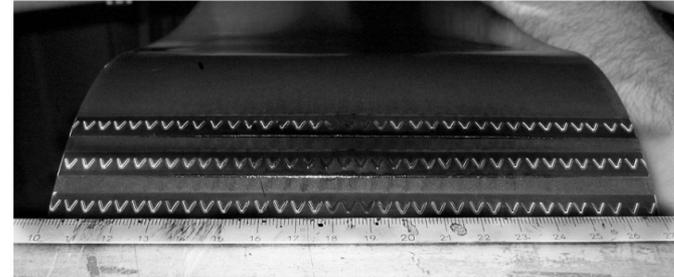
Separation at 1st corner



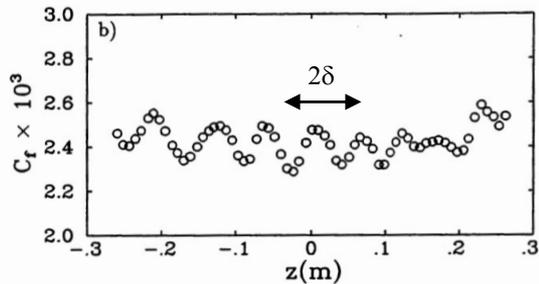
Effective roughness

Upstream history effects

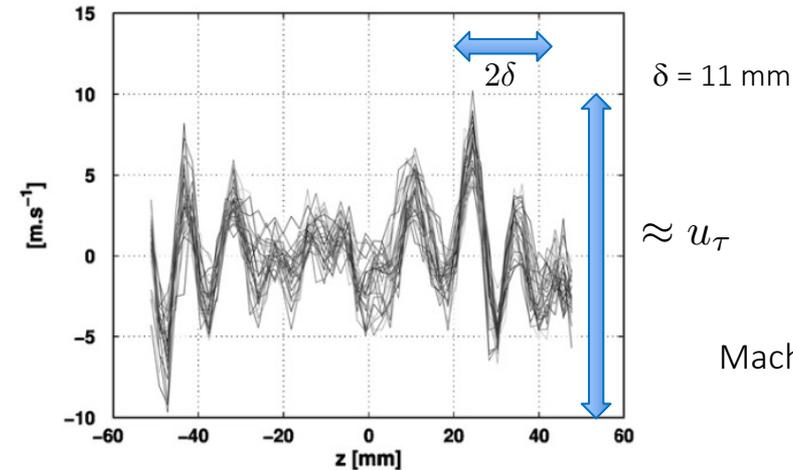
- Periodic roughness upstream of the throat at Mach 2.3
- Vortex generation in subsonic flow
 - Subsonic experiment, following development of streamwise vortices on a wind tunnel wall
 - TG vortices form in the contraction, with $\lambda \approx 2\delta$, persisting into the working section through transition
 - As the boundary layer grows, the vortices appear to pair so as to maintain $\lambda \approx 2\delta$, so their number decreases with downstream distance



$d = 5 \text{ mm}$



Smith (1993)
 $Re_\tau = 4100$



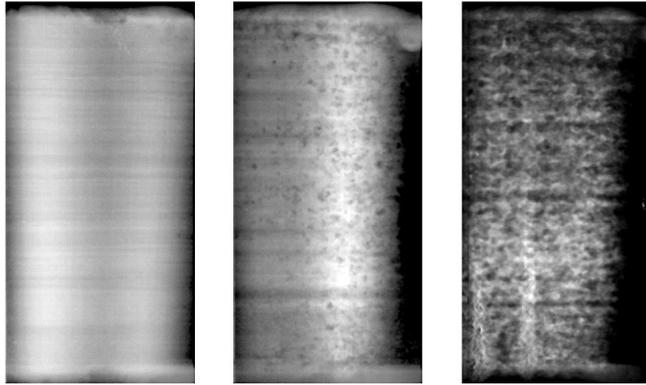
Mach 2.3

Dussauge & Piponnier (2008)

Transitional structures in boundary layers

No tripping

➔ $M = 5$, upstream boundary layer



Laminar

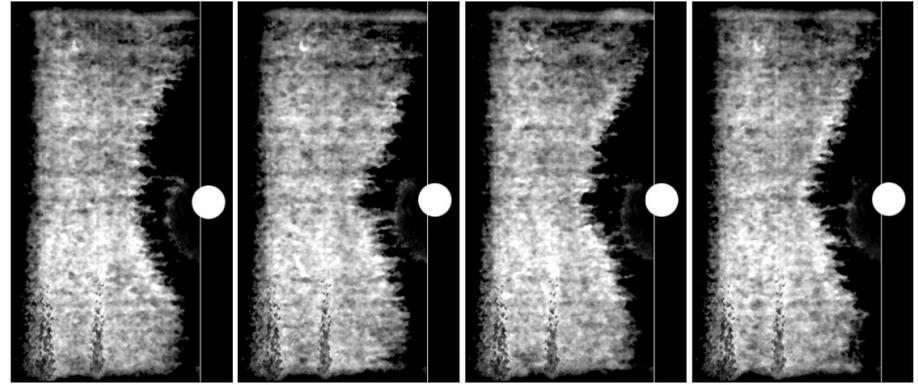
Transitional

Turbulent

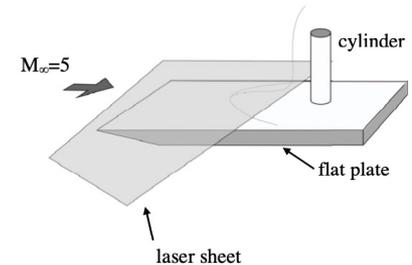
- Organized, longitudinal structures in turbulent boundary layer appear to originate at specific spanwise locations
- Long-lasting

No tripping

➔ $M = 5$, upstream of SWBLI

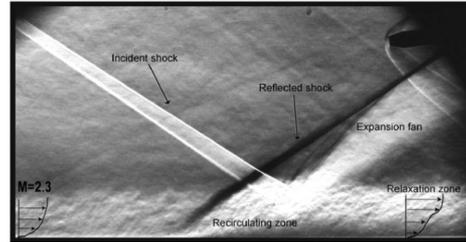
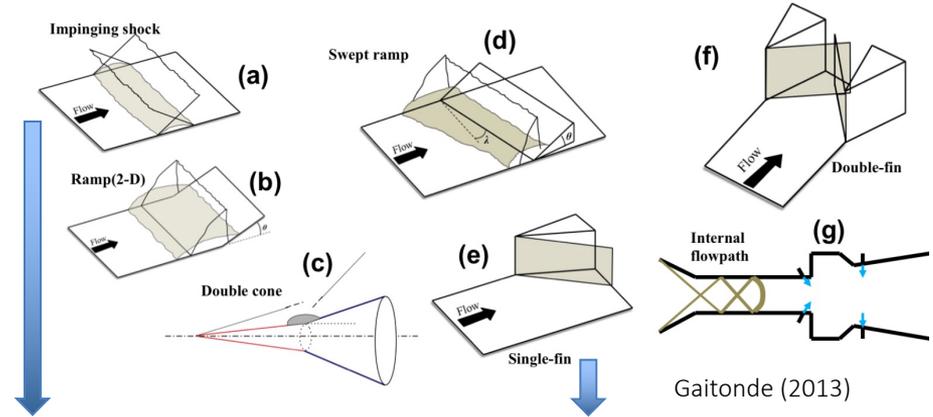


- Organized, longitudinal structures ripple the separation line ahead of cylinder interaction



Shock-wave boundary layer interactions (SWBLI)

- Pressure gradient
- Shock-induced separation
- Streamline curvature effects
- Transverse curvature effects
- Upstream history effects, initial conditions
- Reynolds number effects
- Unsteadiness
- Three-dimensionality
- Asymmetry



Dupont et al. (2005) 9.5° reflected shock at Mach 2.3

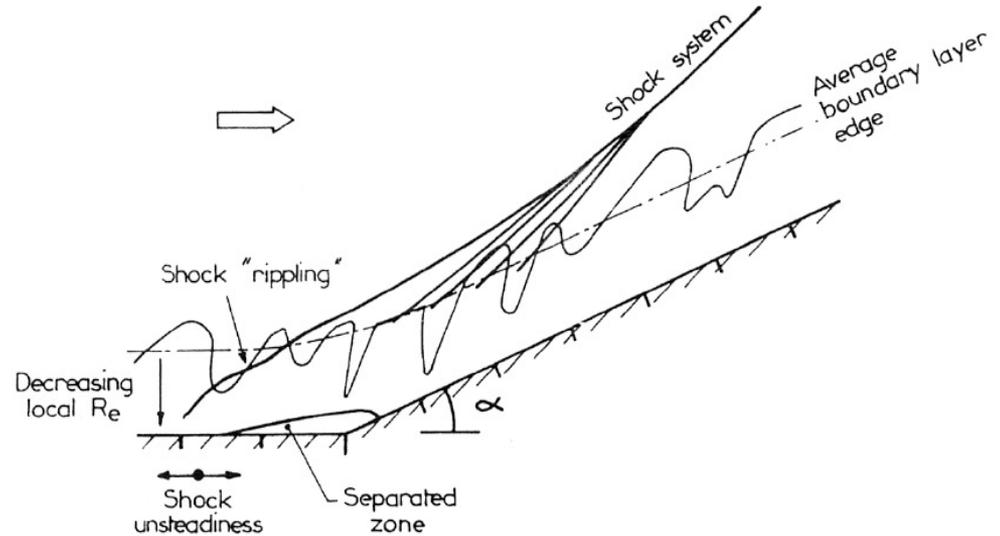


Bookey et al. (2001) 10° sharp fin at Mach 8

Dolling 2001; Délerly & Dussauge 2009; Babinsky & Harvey 2011; Gaitonde 2015

Shock-wave boundary layer interactions (SWBLI)

- Intense shock unsteadiness
- Shock splits and ripples
- Strong shocks lead to separation
- Shock interacts with boundary layer to amplify turbulence
- Peaks occur in friction and heat transfer



Unsteadiness of SWBLI and superstructures

- Shock motion in SWBLI increases with shock strength
- Can lead to severe unsteady pressure and heat loading
- In unseparated flows, VLSM + TG vortices
- In separated flows, VLSM + shear layer instability + TG vortices

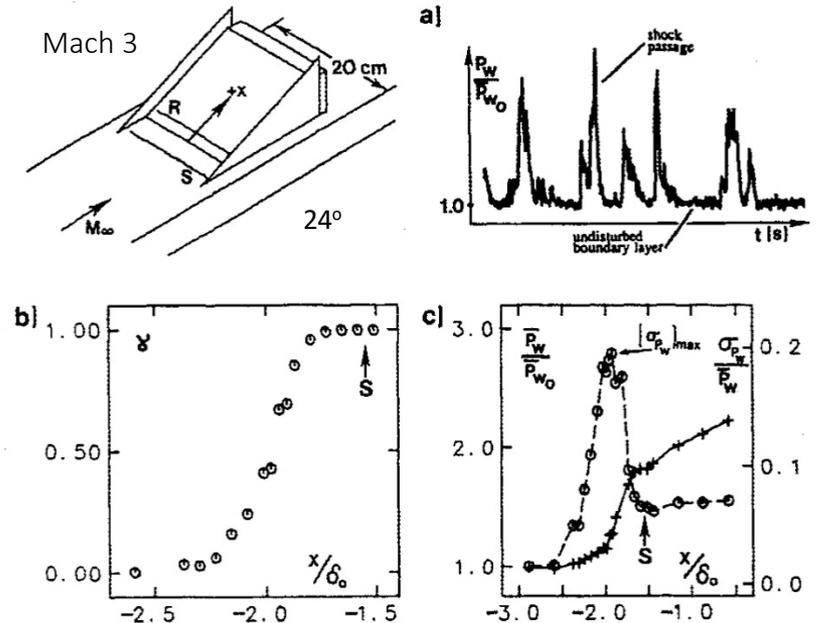


Fig. 1 Typical separated compression ramp flowfield surface properties: a) pressure signal near separation; b) intermittency; and c) mean wall pressure and standard deviation distributions.³

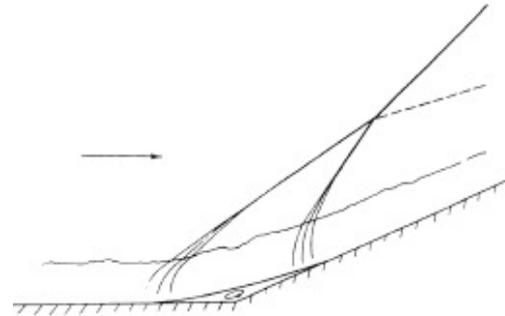
SWBLI at Mach 2.5

500 KHz sequential images of using
FRS (Miles et al.)

MHz rate pulse-burst laser,
coupled with MHz rate
CCD framing camera (PSI)

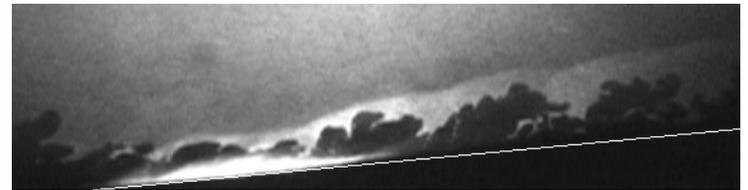
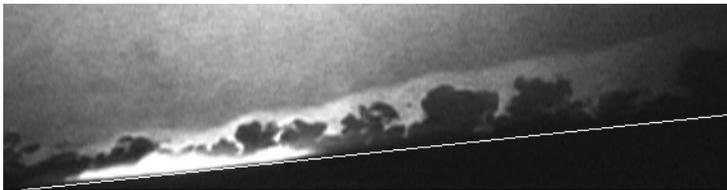
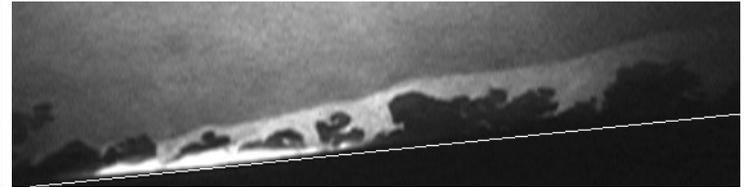
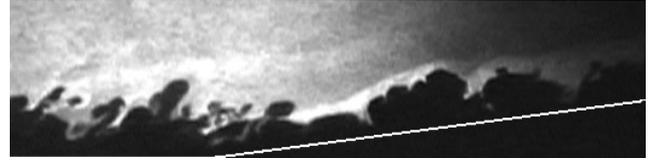
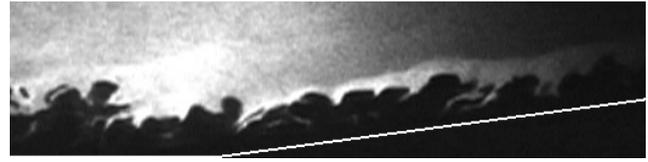
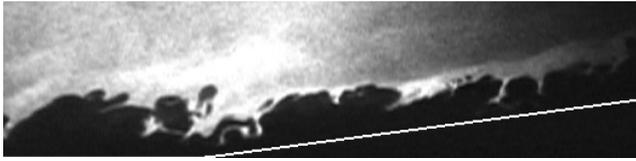
side view

plan view

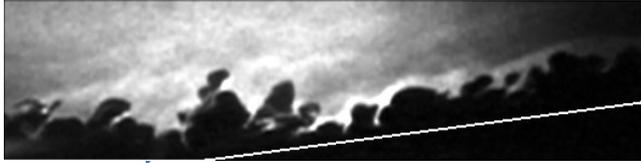


FRS visualizations 8° compression corner at Ma = 7.2

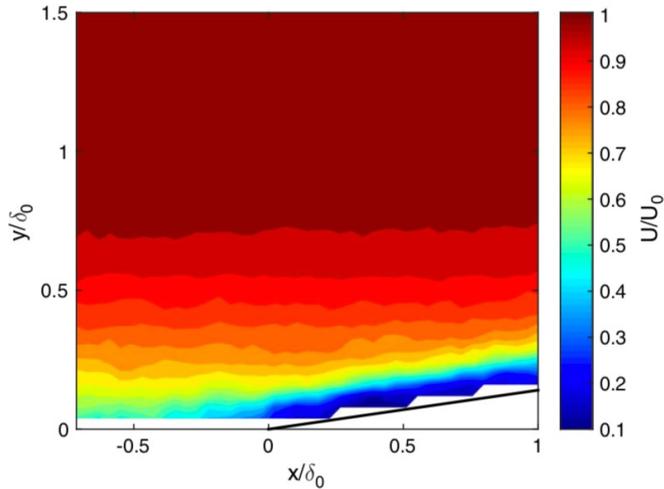
Filtered
Rayleigh
Scattering
Images are
6.5 x 1.8 δ
($\delta = 11$ mm)



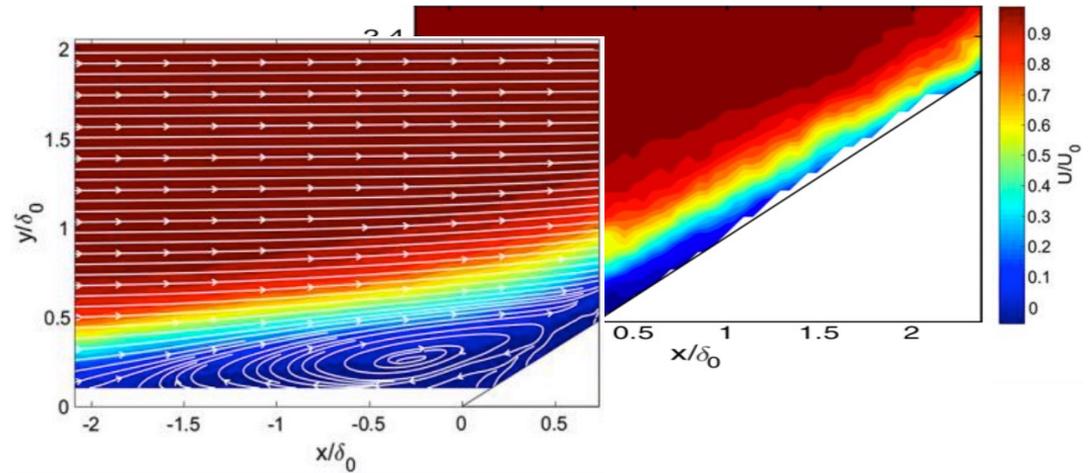
PIV data for 8° and 33° compression corners at $Ma = 7.2$



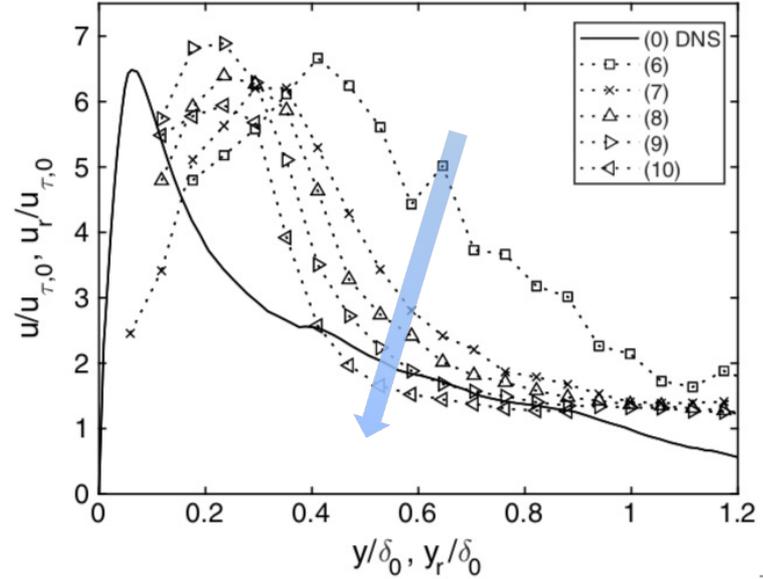
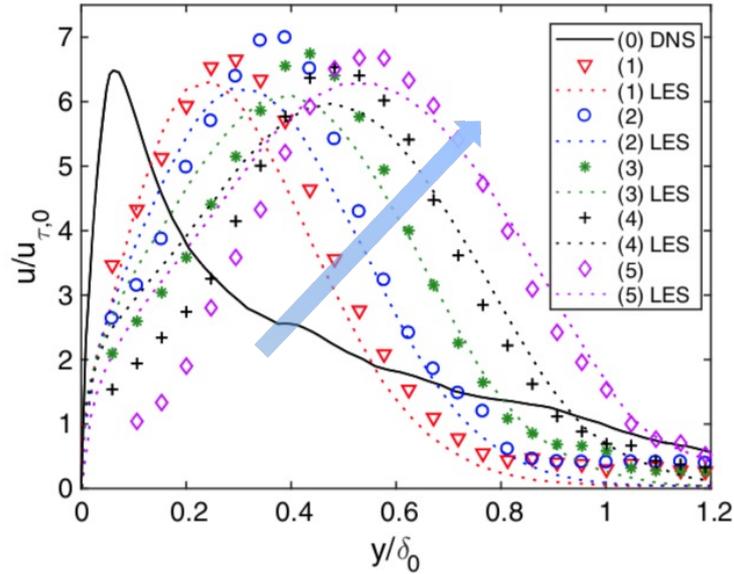
8° corner



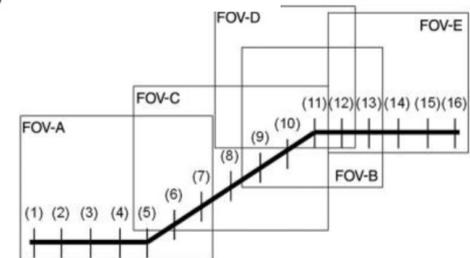
33° corner



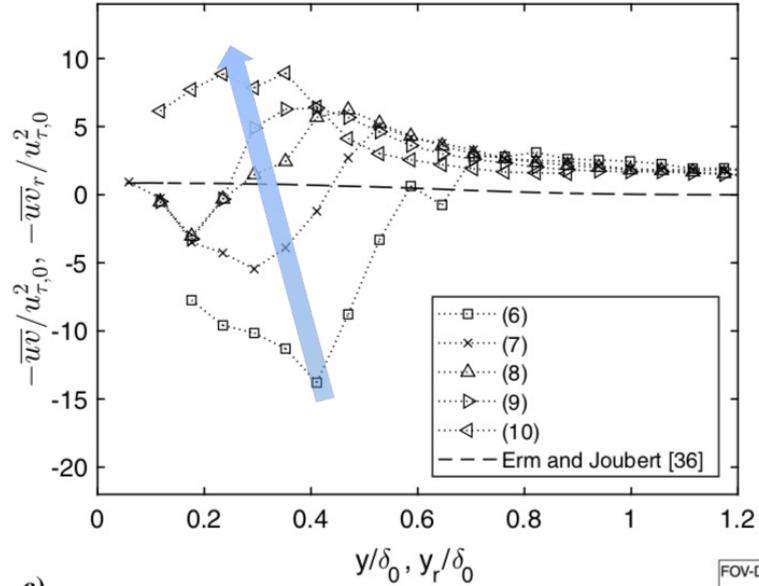
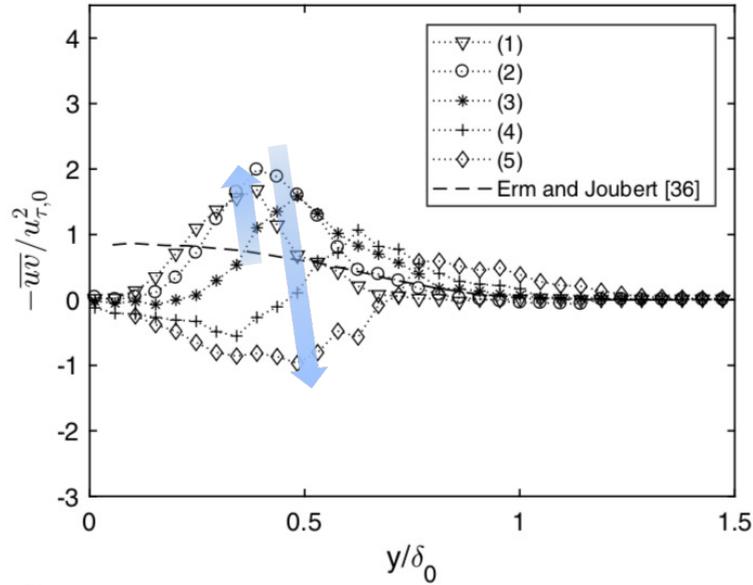
PIV data for 33° compression corner at Ma = 7.2



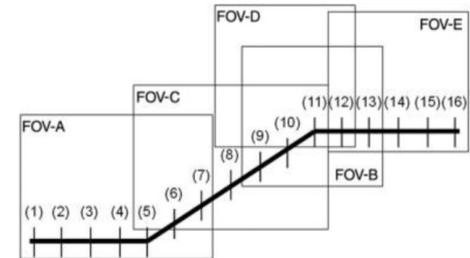
- Progressively larger amplification through the shock
- Morkovin scaling no longer works



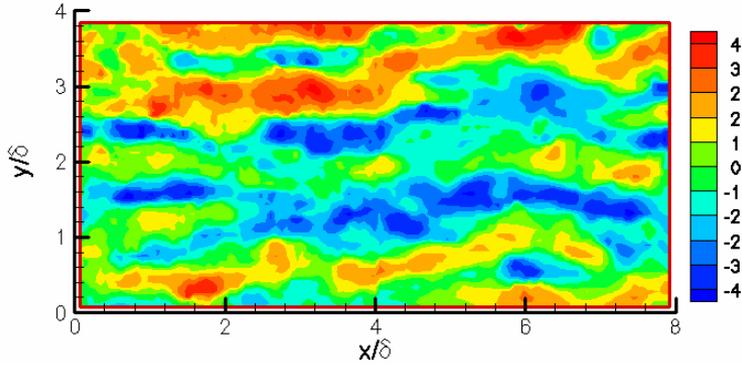
PIV data for 33° compression corner at Ma = 7.2



- Progressively larger amplification through the shock
- Morkovin scaling no longer works

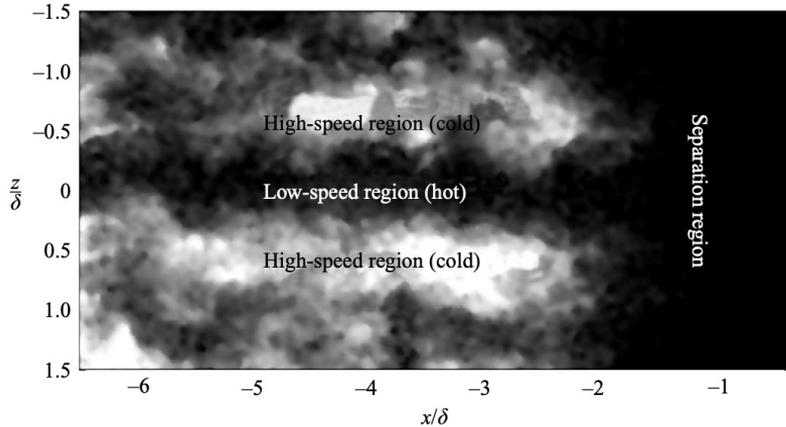
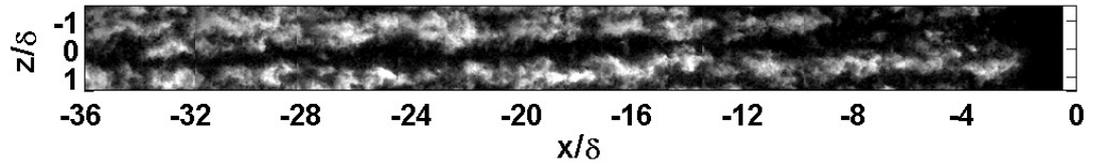


Superstructures in supersonic flows



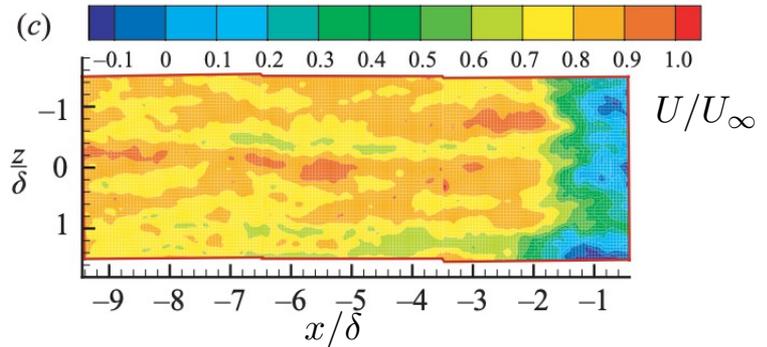
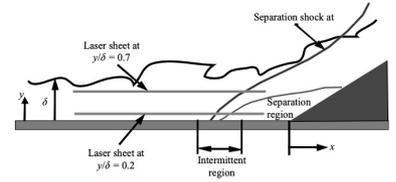
Mach 2 flat plate boundary layer

Ganapathisubramani, et al. (2006; 2007)

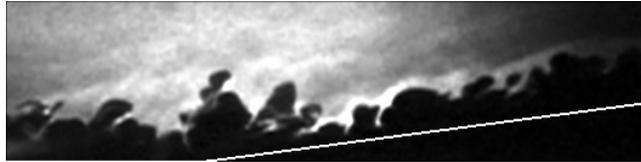


Mach 2 compression ramp 20°

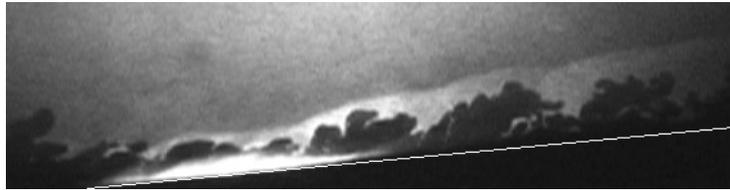
Ganapathisubramani, et al. (2007)



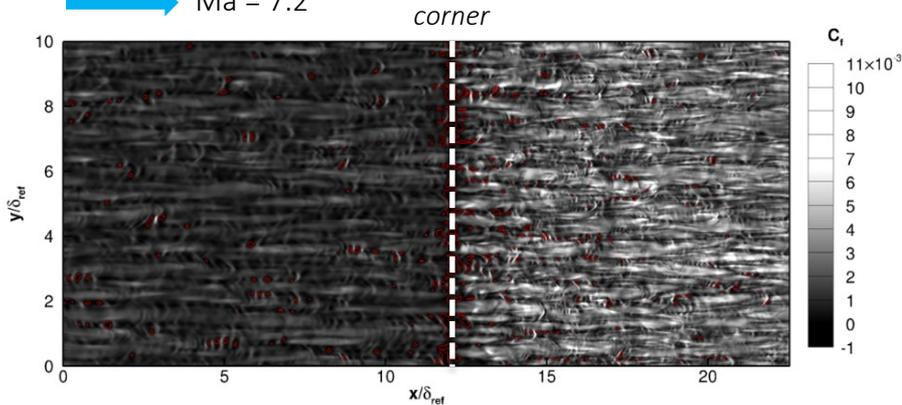
Superstructures and SWBLI



$Re_\tau = 400$
FRS

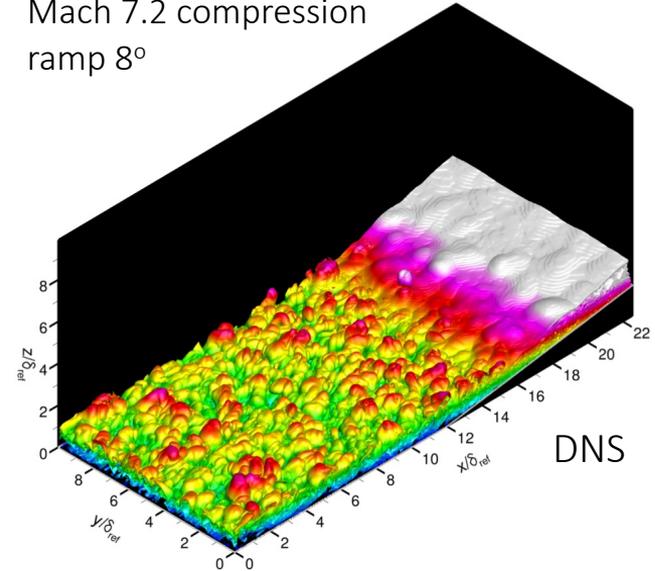


Ma = 7.2



$Re_\tau = 210$

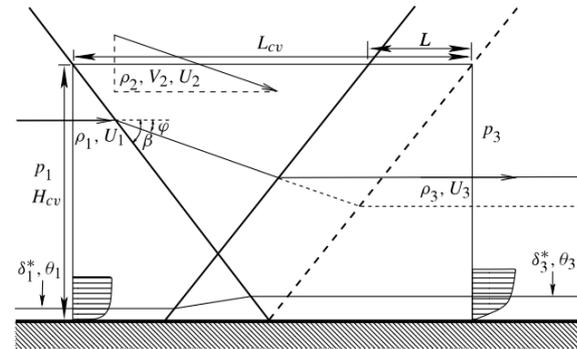
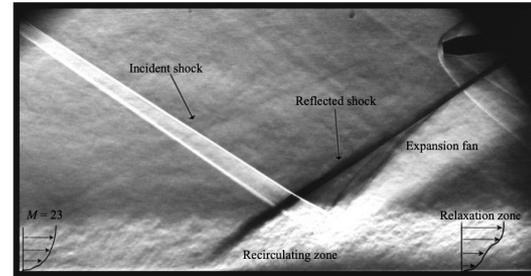
Mach 7.2 compression
ramp 8°



Superstructures modulate the shock and
cause unsteadiness (unseparated flows)

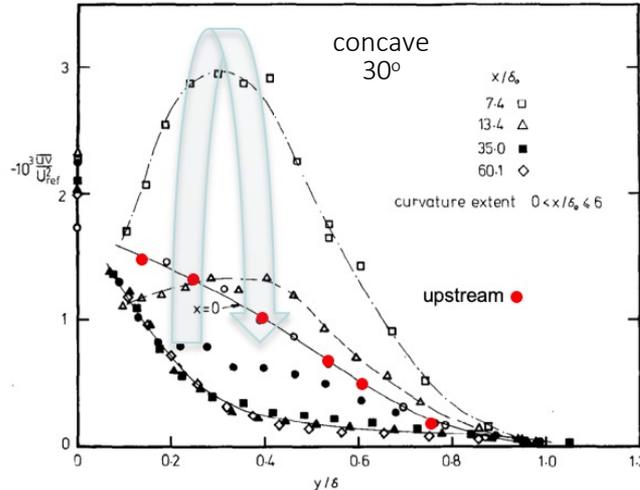
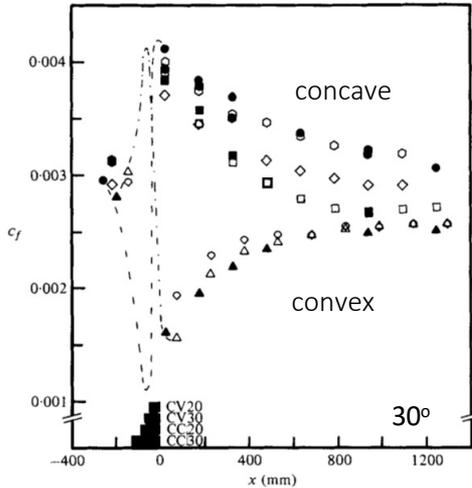
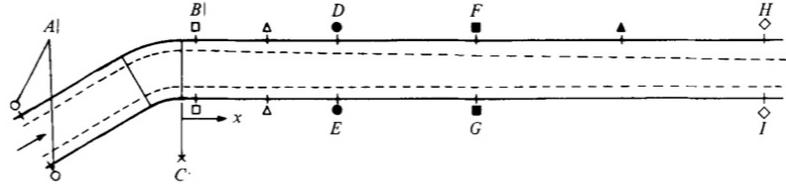
Unsteadiness of separated SWBLI

- Shock motion and unsteadiness in SWBLI increases with shock strength
- In unseparated flows, the superstructures appear to be the prime cause of the shock motion (relatively high frequencies)
- In separated flows, there is a large-scale, slow pulsation of the separation bubble (relatively low frequencies)
- Pulsation related to the entrainment of fluid into the mixing layer feeding the separation bubble
- Leads to a mass balance model of SWBLI interactions and a separation criterion based on Ma , Re , and turning angle

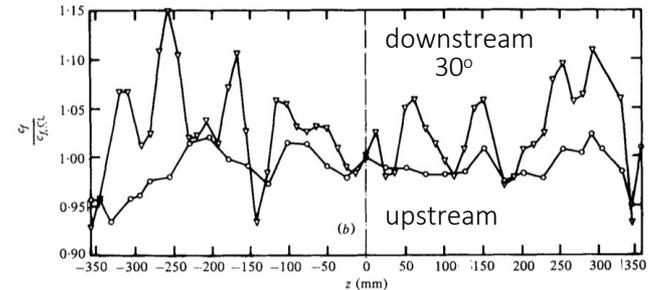


Concave streamline curvature and Görtler vortices

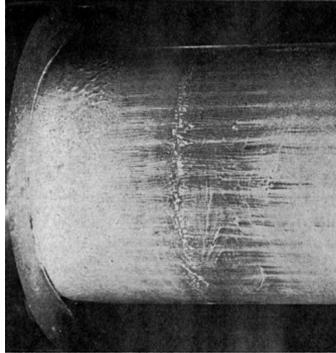
Effect of short regions of curvature on turbulent boundary layers



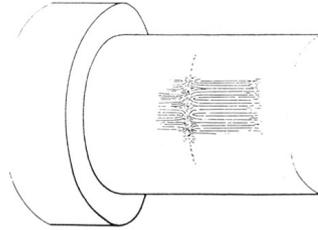
- Convex curvature is stabilizing
 - Expected to promote separation
 - Slow recovery
- Concave curvature is destabilizing
 - Expected to delay separation
 - Non-monotonic flow recovery
 - Appearance of TG vortices



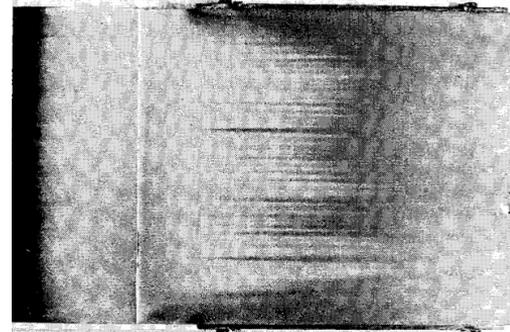
Concave streamline curvature and Görtler vortices



Mach 3.2

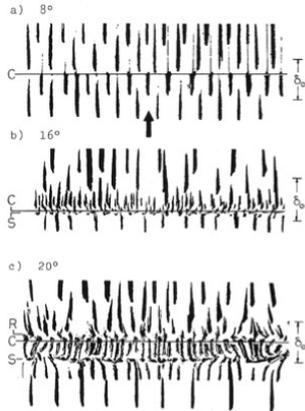


Roshko & Thomke (1966)
Shamroth & MacDonald (1970)

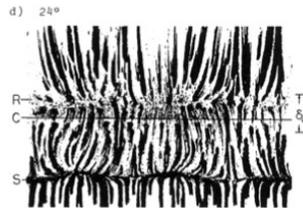


Mach 5.8

Ginoux (1971)
 $\lambda = 2 - 3\delta$



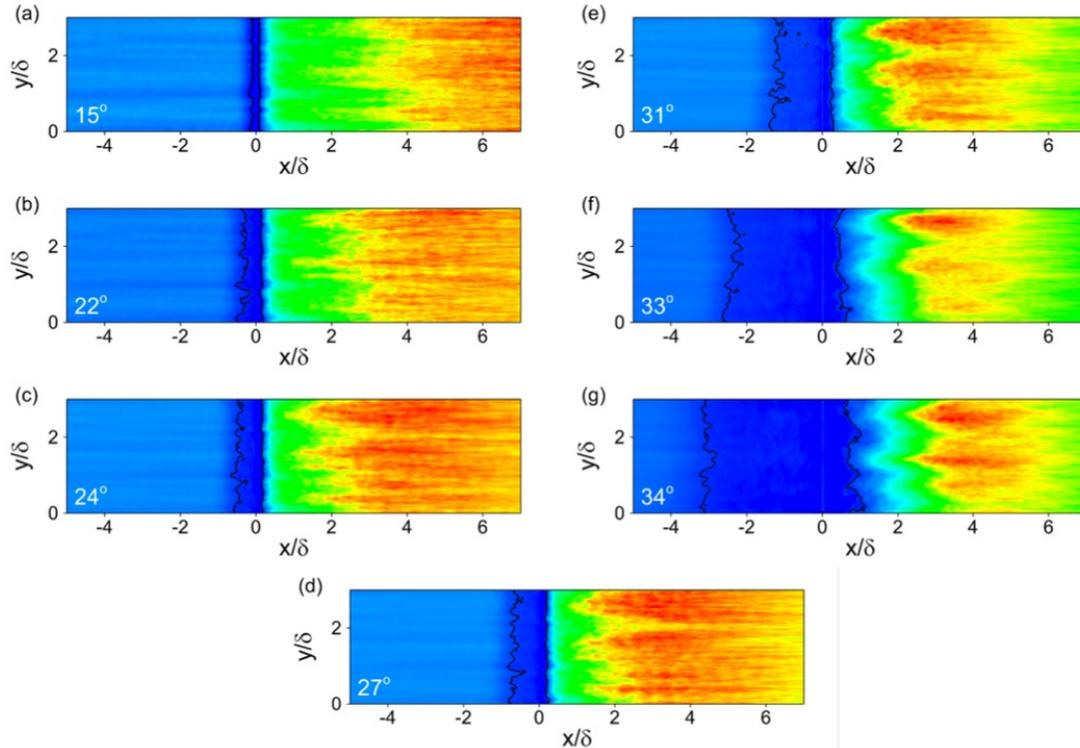
SWBLI Mach 2.9, 8° to 24°
Settles et al. (1979)



“Therefore, although one can conceive that orderly arrays of pairs of counter-rotating vortices of the Gortler type, existing in the free shear layers as a result of streamline curvature, imply this type of flow attachment, one can also conceive that vortices are being developed as a result of this being the only possible type of attachment in the presence of any weak initial disturbance in the approaching boundary layer.” (Ginoux 1971)

Superstructures and separated SWBLI

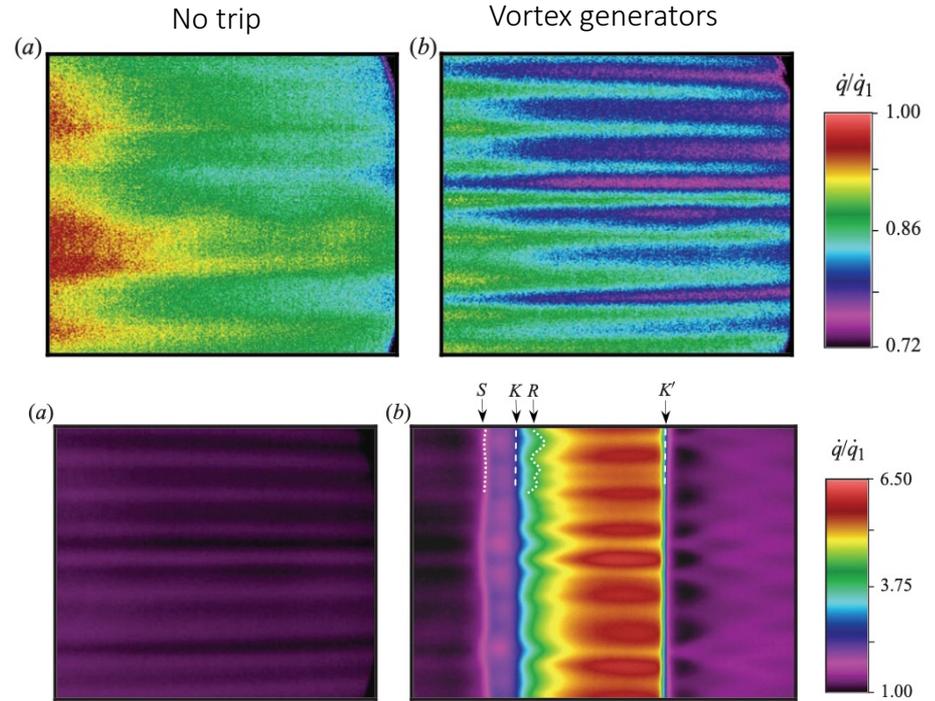
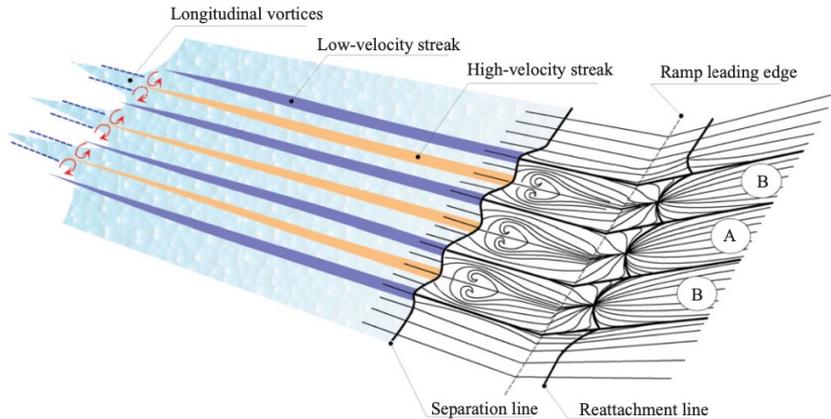
Ma = 10 $Re_\tau = 460$



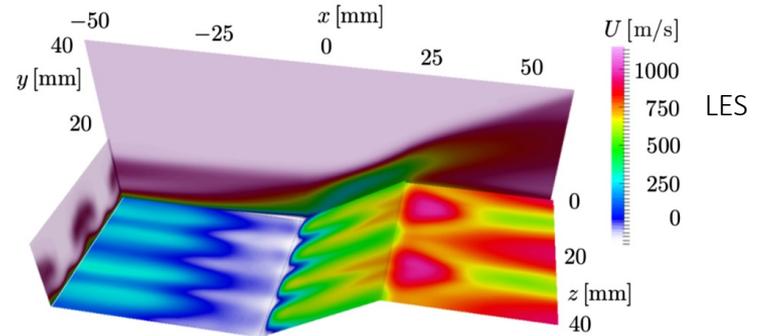
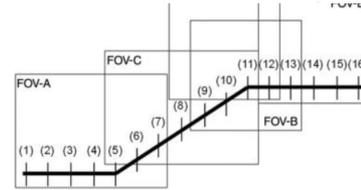
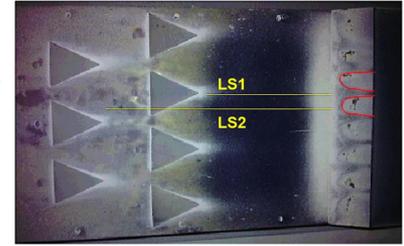
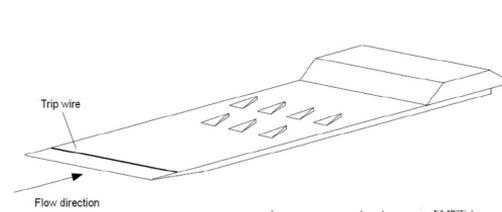
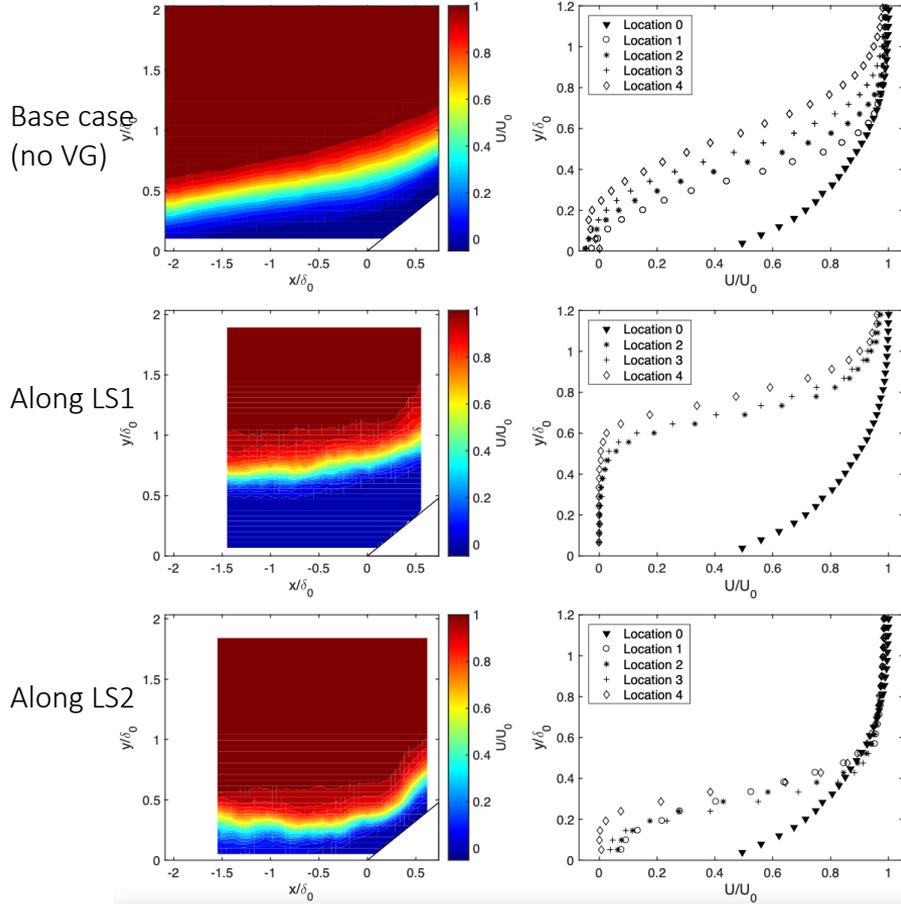
Turbulent boundary layer
compression ramp interactions (LES)

Schulein & Tromifov 2011 (Mach 5)

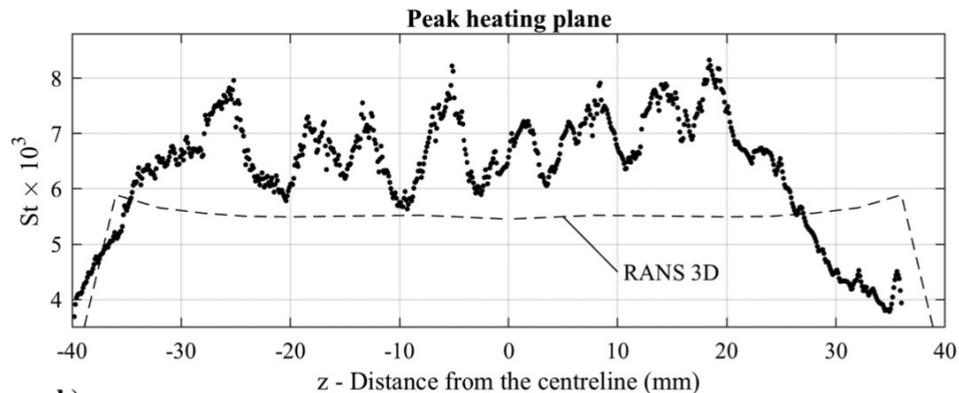
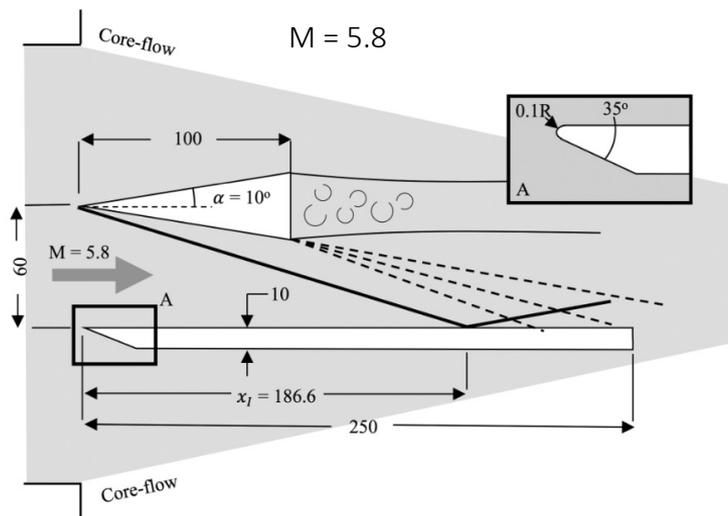
- Longitudinal vortices, initiated at the leading edge, survive for long distances, $O(10^4 h)$
- The total number, dimension and type of installed vortex generators define the number of vortex pairs generated downstream



VG control for 33° compression corner at Ma = 7.2

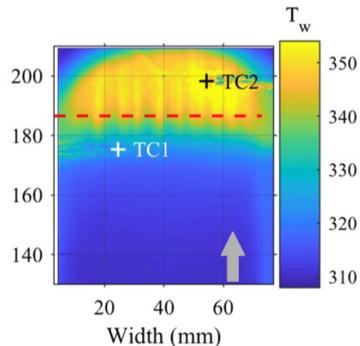


Separation by reflected shock Mach 5.8



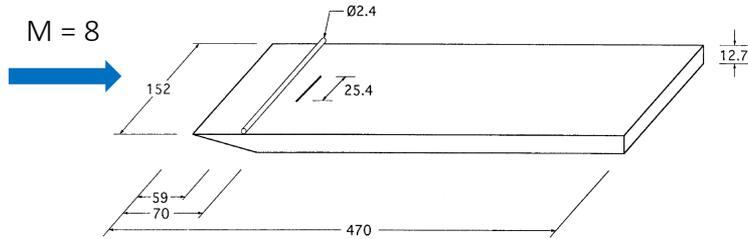
Currao et al. (2020)

Temperature
plan view



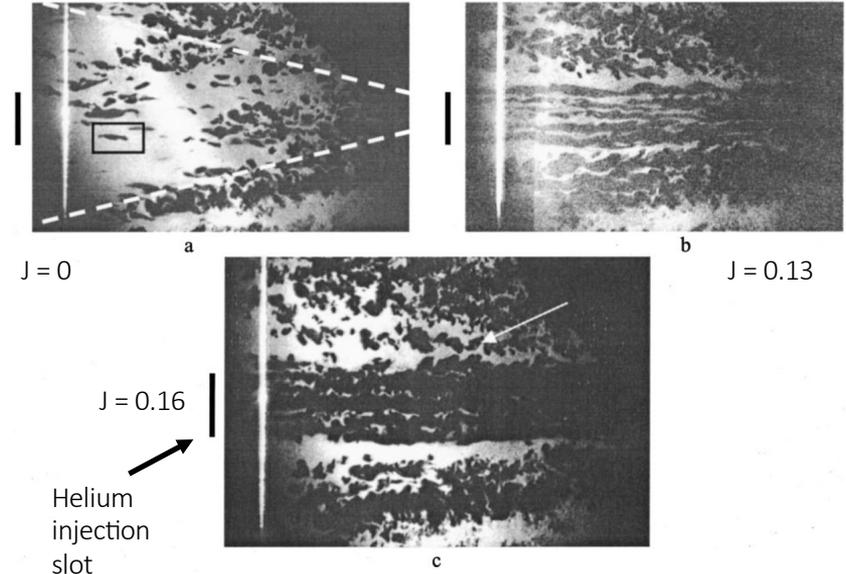
- Boundary layer transitions to turbulent on shock impingement
- Taylor-Görtler vortices form almost immediately
- Heat flux varies by $\pm 15\%$
- Missed by RANS

Helium injection in Mach 8 turbulent boundary layer



$$J = \frac{(\rho U^2)_{inj}}{(\rho U^2)_{\infty}} = \frac{(\gamma p M^2)_{inj}}{(\gamma p M^2)_{\infty}}$$

- Organized, longitudinal structures in turbulent boundary layer are formed in region of helium injection
- Taylor-Görtler vortices due to concave curvature caused by injection



Görtler number

- Taylor-Görtler vortices in laminar, incompressible flows with concave streamline curvature are seen to occur when (Liepmann, Schlichting, Dryden)

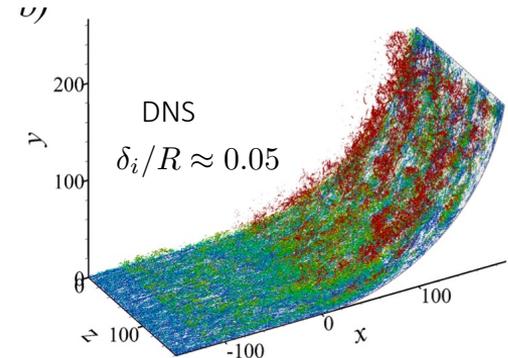
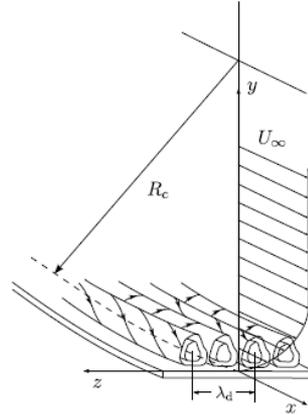
$$G_\theta = Re_\theta \sqrt{\frac{\theta}{R}} > [6, 9]$$

- For turbulent, incompressible flows with concave streamline curvature, Tani (1962) suggested

$$G_\theta = \frac{U_\infty \theta}{\nu_T} \sqrt{\frac{\theta}{R}} = \frac{55}{H} \sqrt{\frac{\theta}{R}} > [\approx 4]$$

with $\nu_T = 0.018 U_\infty \delta^*$ (Clauser)

- For turbulent, compressible flows with concave streamline curvature, Smits & Dussauge (2006) suggested incorporating compressible forms of H and θ
- For a given δ , H increases faster than θ decreases with M, so the Görtler number decreases with M
- Boundary layers more stable to concave curvature still holds with $\nu_T \propto U_\infty \delta_i^*$ (Maise & McDonald)

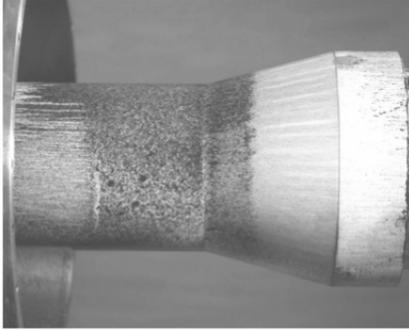


You et al. (2021)

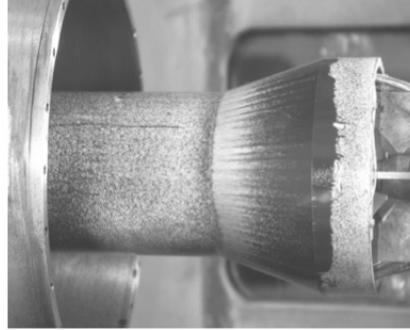
Are these simple ideas correct?

Separation on cylinder/flare Mach 5

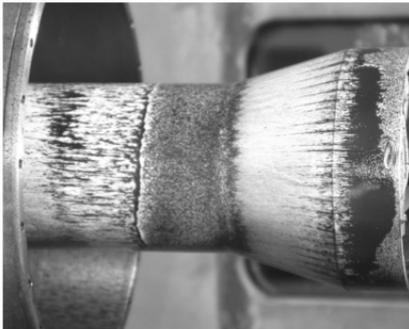
M = 5, axisymmetric flare/cone
Oil flow visualization in natural transition



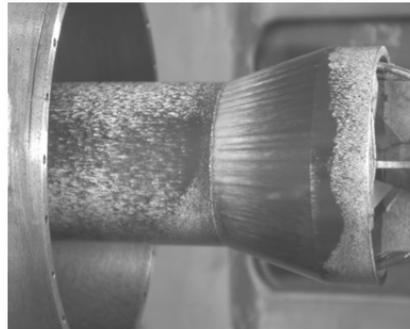
a) $p_{st} = 0.9 \times 10^5$ Pa, $Re_L = 0.38 \times 10^6$



c) $p_{st} = 5.4 \times 10^5$ Pa, $Re_L = 1.46 \times 10^6$



b) $p_{st} = 2.1 \times 10^5$ Pa, $Re_L = 0.67 \times 10^6$



d) $p_{st} = 5.5 \times 10^5$ Pa, $Re_L = 1.55 \times 10^6$

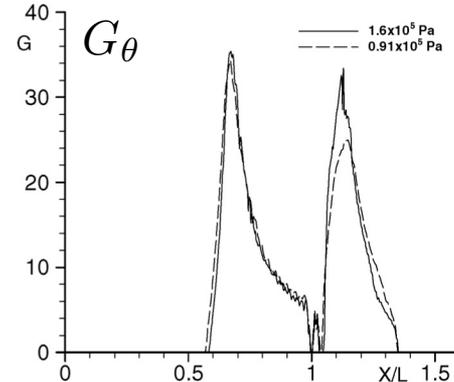
- Taylor-Görtler vortices present in all cases

$$G_\theta = Re_\theta \sqrt{\frac{\theta}{R}}$$

$$G_{\delta^*} = Re_{\delta^*} \sqrt{\frac{\delta^*}{R}} = H^{3/2} G_\theta$$

$$H \approx (1 + 0.693(\gamma - 1)M_\infty^2) \bar{H} = 20.5 \quad \text{at } M_\infty = 5$$

$$G_{\delta^*} \approx 90 G_\theta$$

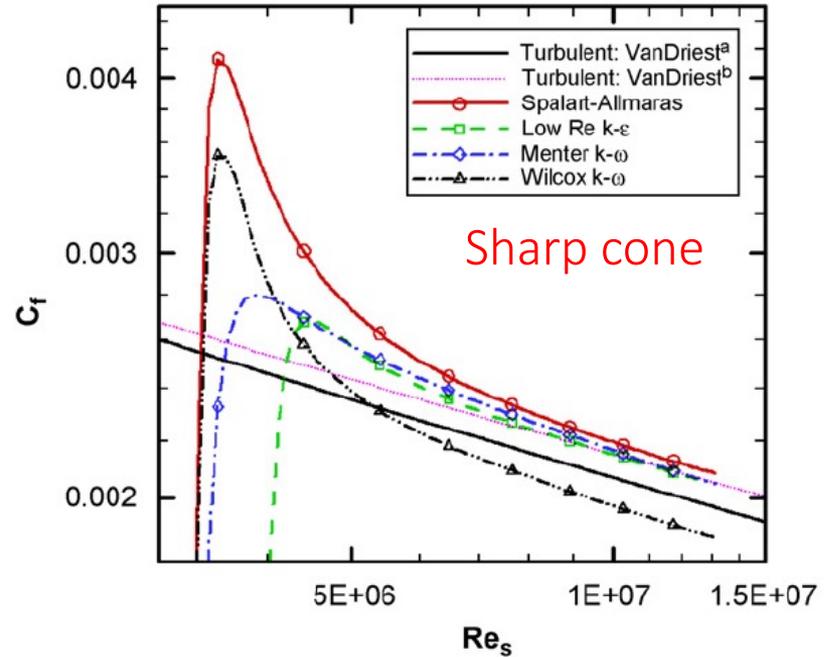
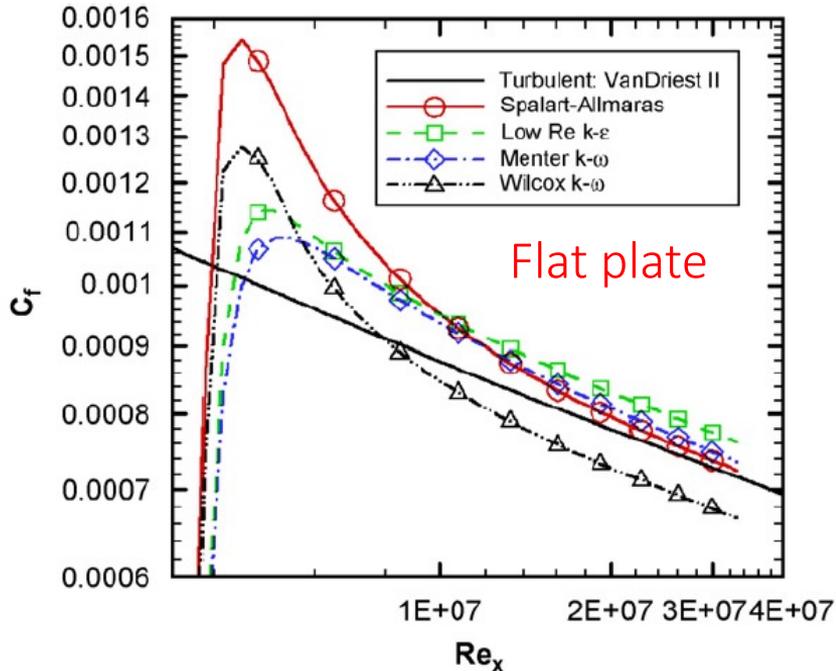


$$Re_L = 0.38 \times 10^6,$$

$$= 0.68 \times 10^6$$

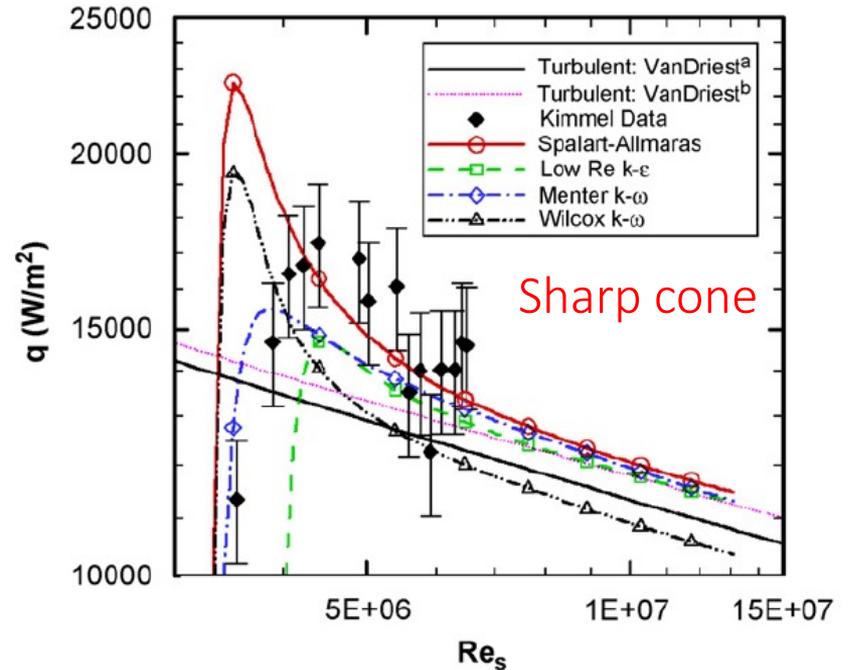
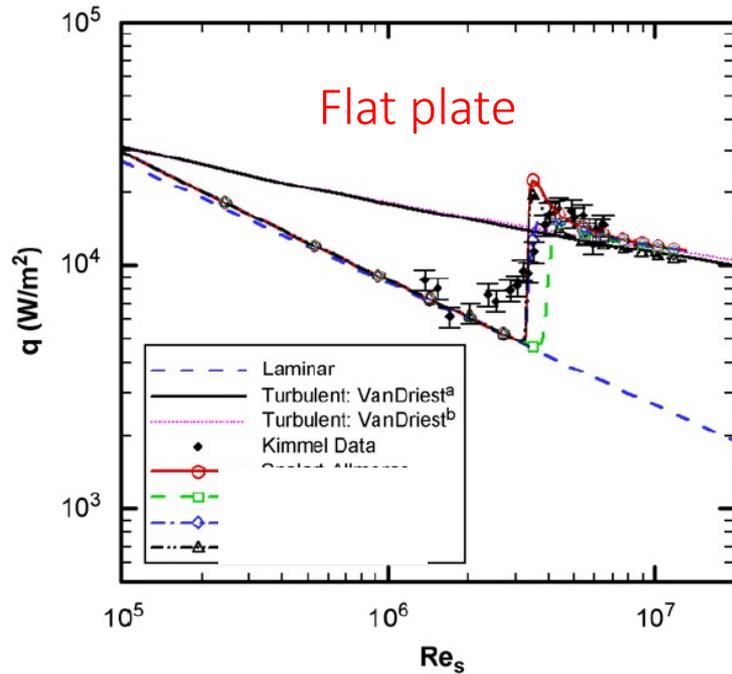
What about turbulence modeling (RANS)?

Skin friction at $Ma = 8$



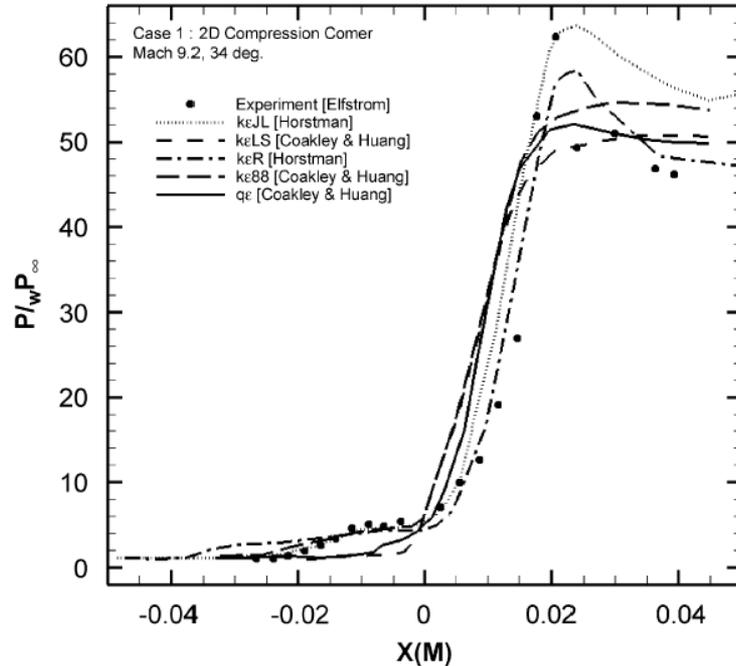
What about turbulence modeling (RANS)?

Heat transfer at $Ma = 8$

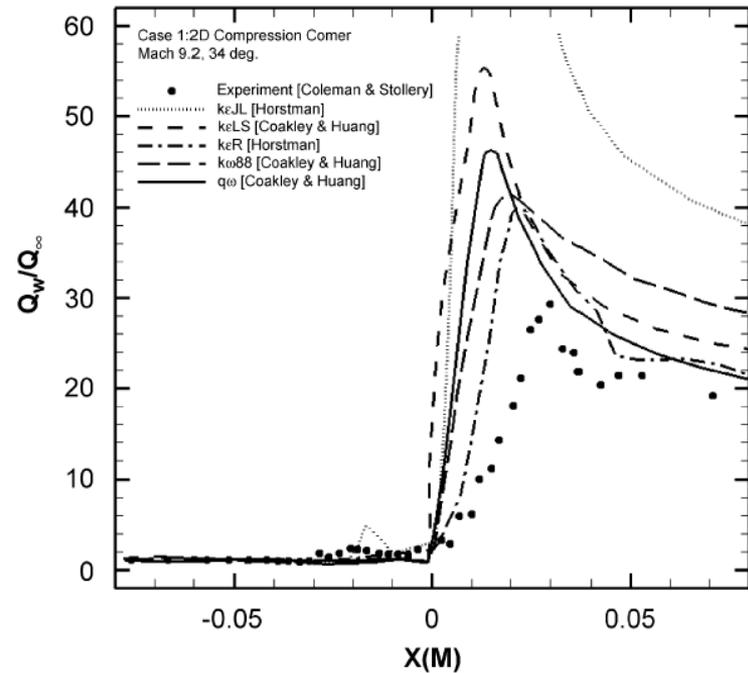


34° compression corner at Ma = 9.2 (separated)

Wall pressure



Heat transfer



Summary

- Zero pressure gradient, smooth wall flows appear to follow Van Driest/Morkovin scaling for mean flow and turbulence
- Shock-wave boundary-layer interactions are inherently unsteady, with large fluctuations in wall pressure and heat transfer
- SWBLI can cause very large amplifications of turbulence
- RANS models not reliable for SWBLI, especially at high Mach number

Remarks

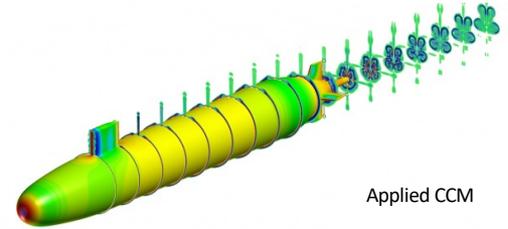
- Increasing Mach number generally coupled with decreasing Reynolds number
 - perfect for DNS (or high-resolution LES)
- High Mach number experiments must be designed for DNS
- Experiment and DNS/LES must be done at the same Reynolds number
- Need to expand our interests to include
 - heat transfer, roughness, film cooling
 - ablation, wall catalysis, pyrolysis, real gas effects, chemistry
 - curvature, pressure gradient, three-dimensionality
- Diagnostics are crucial to everything (understanding, modeling, validation, etc.), especially for turbulence
 - New diagnostics for τ_w' , q' , species, reaction rate?

Opportunities

- Flows with pressure gradients
 - What constitutes a canonical experiment?
 - Is there one or more characteristic parameters?
 - Prediction of separation on smoothly varying surfaces is still a major challenge
 - Data base is scattered, not systematic nor complete
- Turbulence Modelling
 - LES for separated flows well-established
 - RANS-DES easy for fixed separation point, hard for smoothly varying surfaces
 - LES wall functions necessary to make progress – lots of work needed
 - RANS for compressible flows, especially SWBLI (including transonic flows with possible separation)
 - Hypersonic flows
- Reduced order modeling
 - Resolvent analysis (McKeon)
 - Linear models (Gayme)
- Machine learning/AI
 - Can only be effective if there is a large-enough database

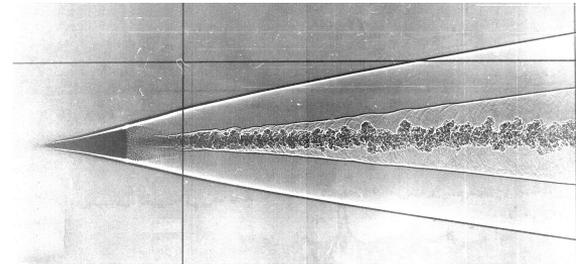
Incompressible turbulence

- Canonical turbulent flows
 - Grid turbulence
 - Jets and wakes
 - Fully-developed flows (channel, pipe, Couette, Ekman, Taylor-Couette)
 - Flat plate boundary layers, smooth or rough walls
- Perturbations in boundary conditions
 - Roughness, heat transfer
 - Step changes, impulsive changes
- Pressure gradients
- Complex flows
 - Streamline curvature (convex and concave)
 - Streamline convergence and divergence
 - Rotation (turbulent vortex flows)
 - Stratification (stable or unstable)
- Bodies of revolution (concurrent complex flows)
- Separation
 - Sudden (backward facing step, sharp edges)
 - Gradual (slowly varying flows)
- Flows with chemistry (mixing)
- Noise generation and transmission
- Flow control

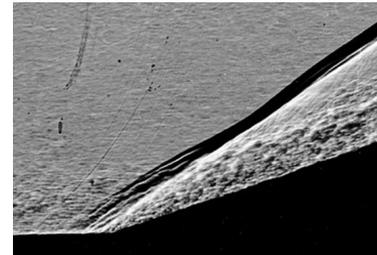


Compressible turbulence

- Canonical turbulent flows
 - Grid turbulence
 - Jets and wakes
 - Channel, pipe
 - Flat plate boundary layers, smooth or rough walls
- Perturbations in boundary conditions
 - Roughness, heat transfer
 - Step changes, impulsive changes
- Pressure gradients
- Complex flows
 - Streamline curvature (convex and concave)
 - Streamline convergence and divergence
 - Rotation (turbulent vortex flows)
 - Stratification (stable or unstable)
- Separation
 - Sudden (backward facing step, sharp edges)
 - Gradual (slowly varying flows)
- Shock-wave boundary layer interactions (2D and 3D)
- Flows with chemistry
- Noise generation and transmission

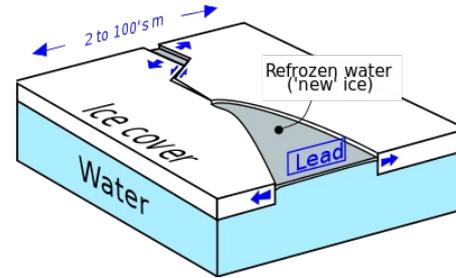


Mach 5.6 (Canning, NASA-Ames)



Some opportunities in turbulence

- Scalar transport
 - Contaminants, food or prey, pheromones
- Two-phase flows
 - Particle/bubble tracking (Lagrangian dynamics)
 - Cloud formation, plankton upwelling
 - Sediment, dust transport
- Atmospheric/geophysical flows
 - Stable (night time), unstable (convective, day time)
 - Rotating flows (Coriolis)
 - Heat, mass and momentum fluxes
 - Step changes (leads and polynyas)
- Stratified turbulence



Experiment, DNS, LES

Golden age of turbulence

Rapidly improving experiment and computation
Democratization of experiment and computation
New facilities, new ideas
Need courage to move beyond canonical flows

Questions?

